

# DO SUPERMASSIVE BLACK HOLES IMPACT A GALAXY'S ABILITY TO FORM NEW STARS?

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AIP NSW Postgraduate Awards, November 17th 2015



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Department of Industry and Science



# WHAT **SUPPRESSES** STAR-FORMATION IN GALAXIES?



**Younger Galaxies**

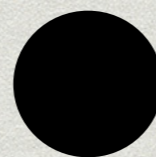
**Transitional**

**Older Galaxies**

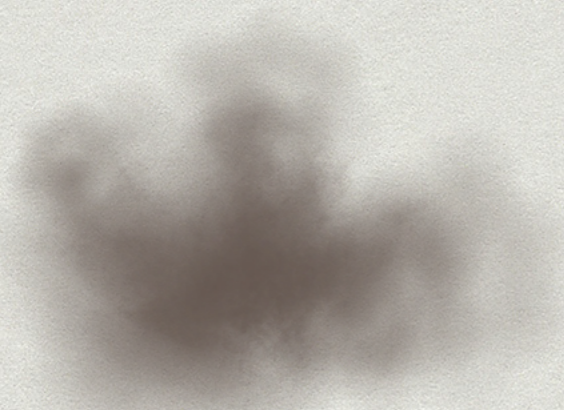
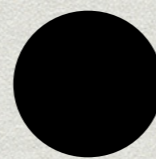
Understanding this is central to building a  
coherent picture of galaxy evolution

Image Credit: NASA/JPL-Caltech

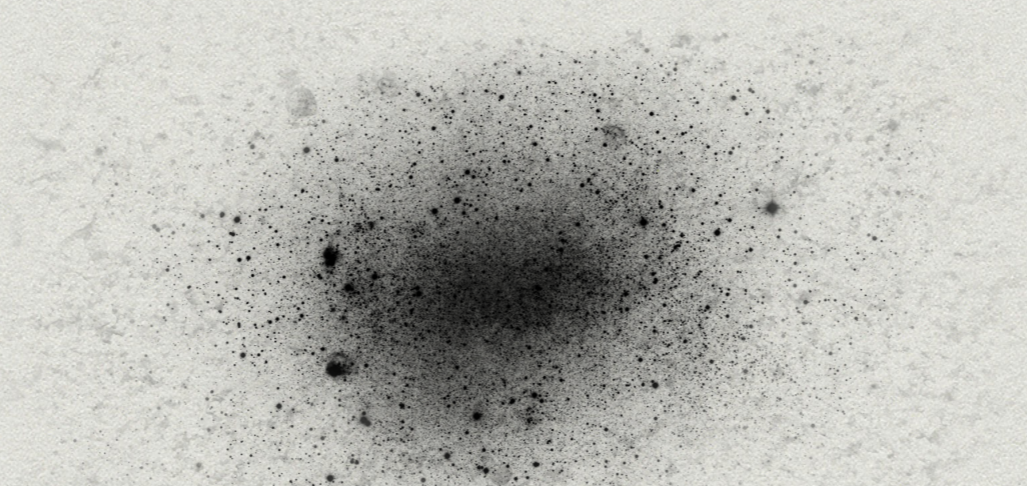
# SUPERMASSIVE BLACK HOLE



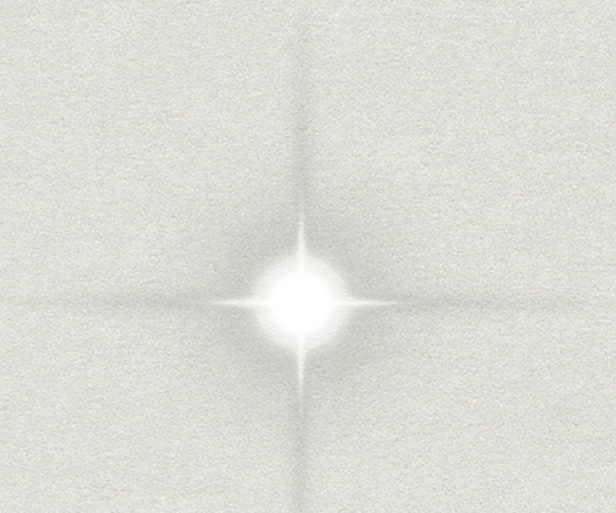
# SUPERMASSIVE BLACK HOLE



gas



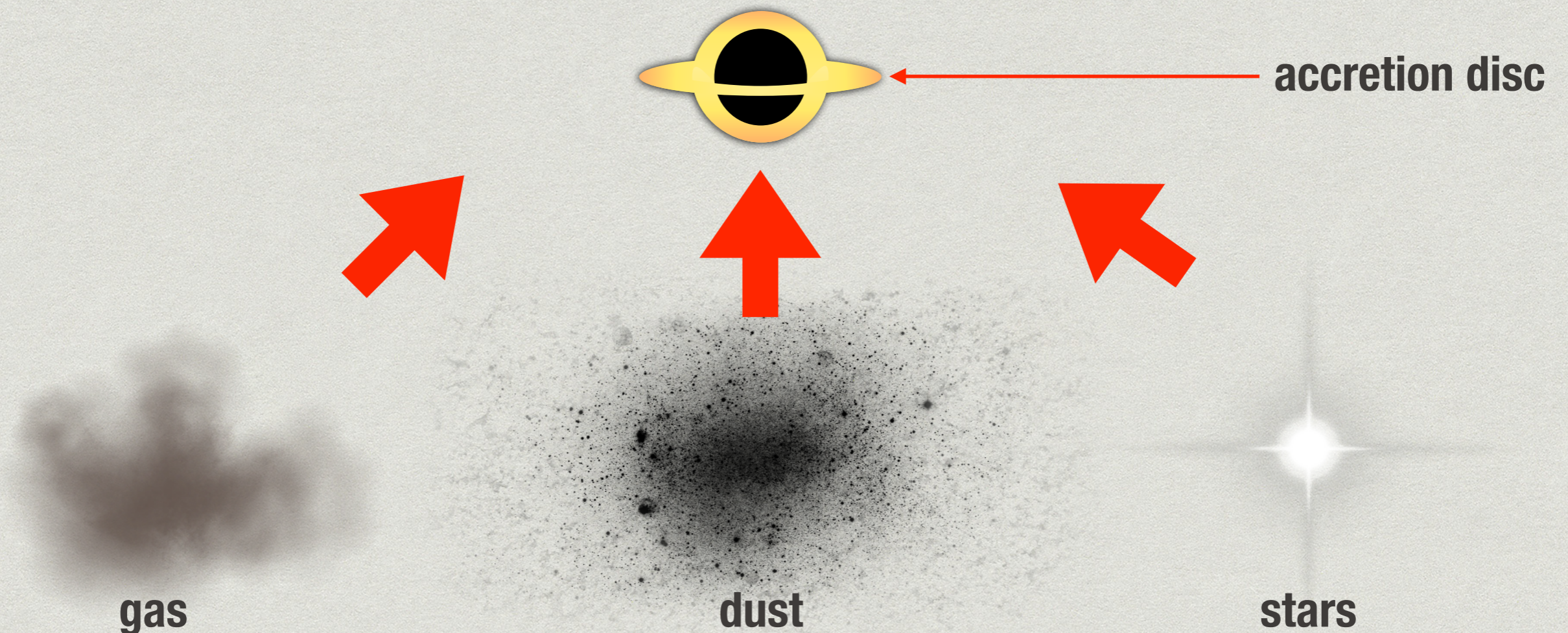
dust



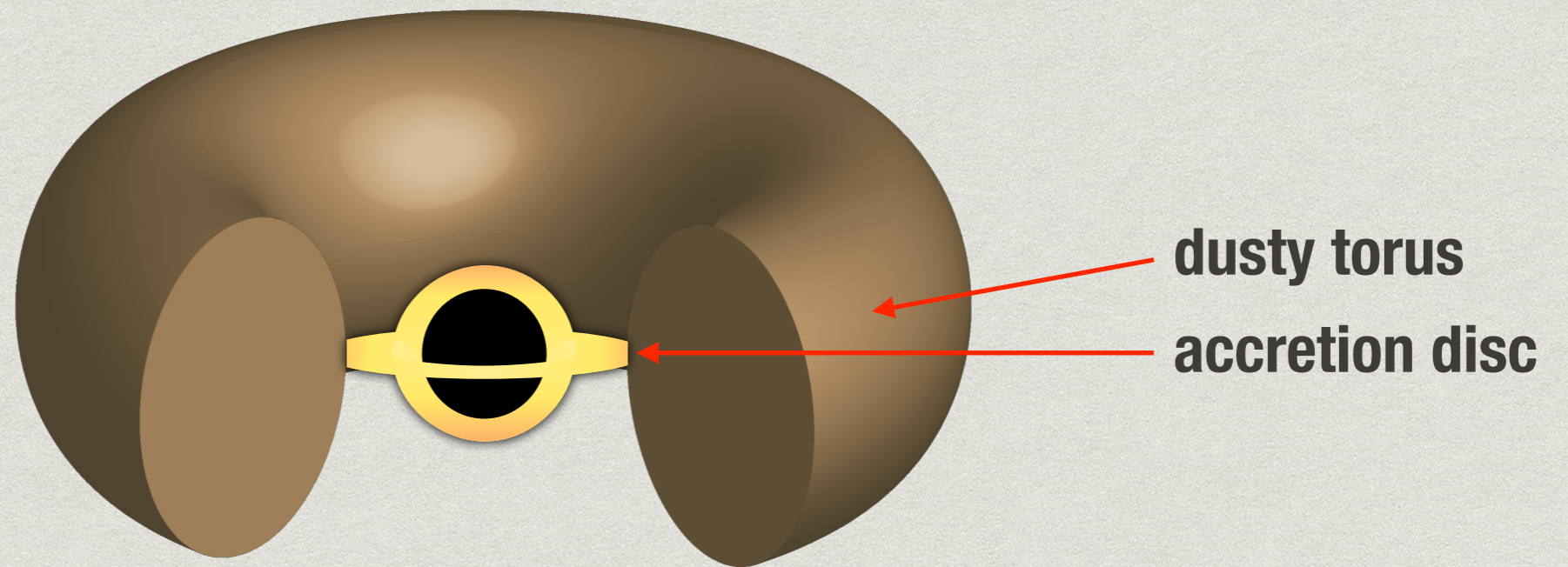
stars

# FEEDING A SUPERMASSIVE BLACK HOLE

Conservation of angular momentum  
causes fuel to form a disk as it spirals in



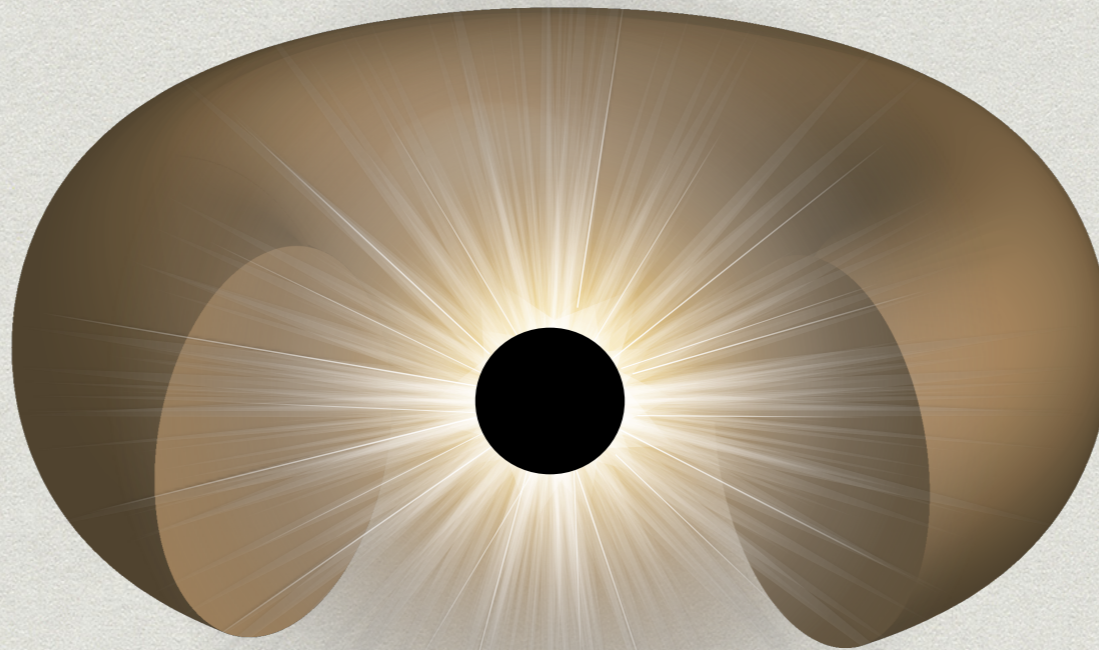
# FEEDING A SUPERMASSIVE BLACK HOLE



# ACTIVE SUPERMASSIVE BLACK HOLE

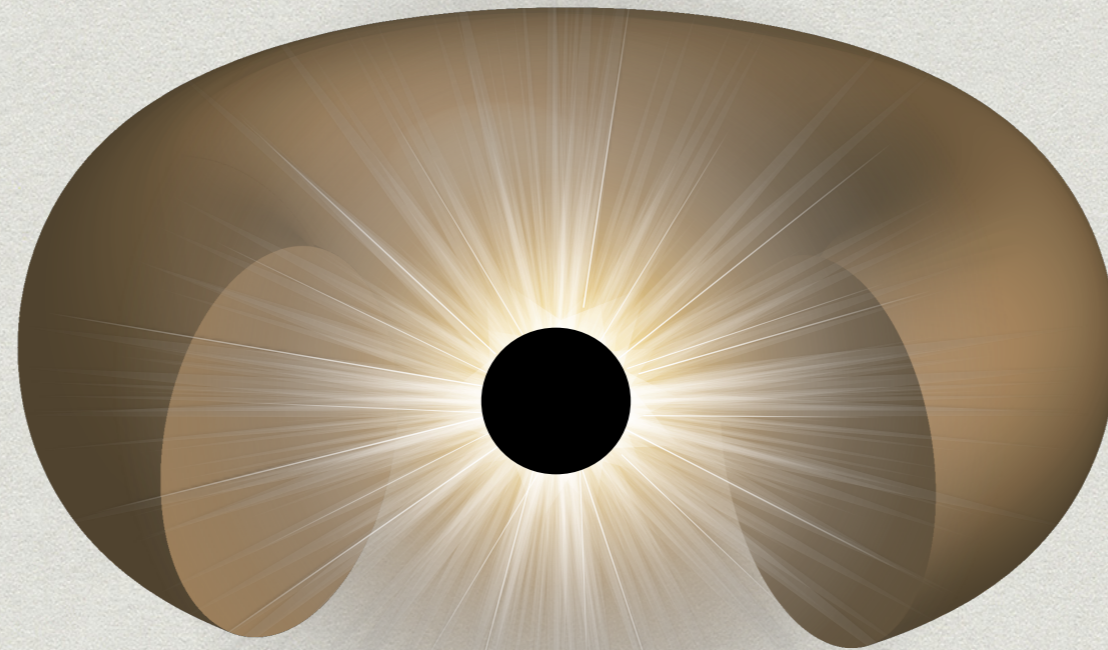
$$L \sim 10^{37-41} \text{ W}$$

(i.e.  $\sim 10^{10-14} L_{\odot}$ )



**Within this disc, there are dissipative processes, e.g. collisions, shocks, etc.  
This dissipated energy emerges as radiation**

# ACTIVE SUPERMASSIVE BLACK HOLE



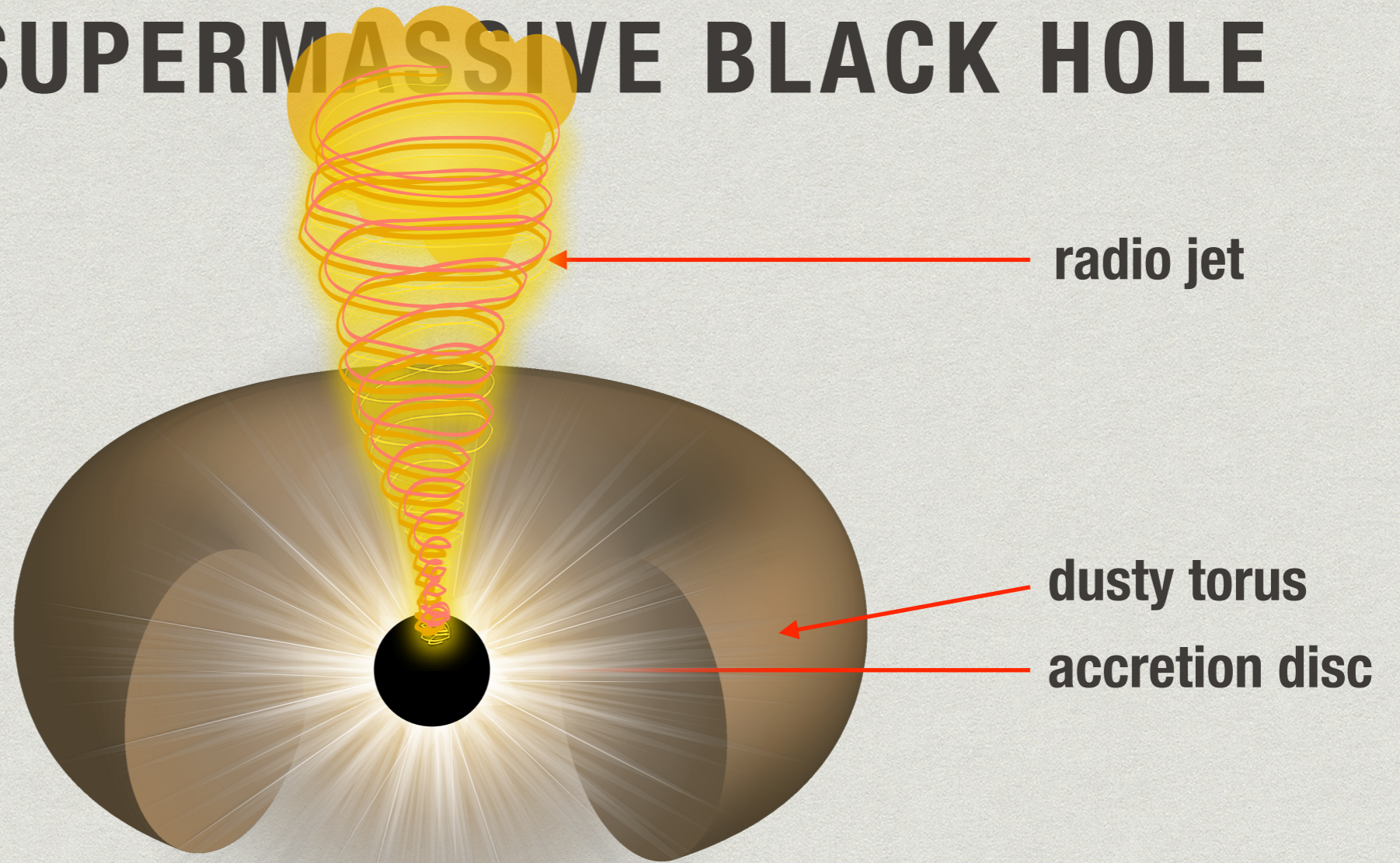
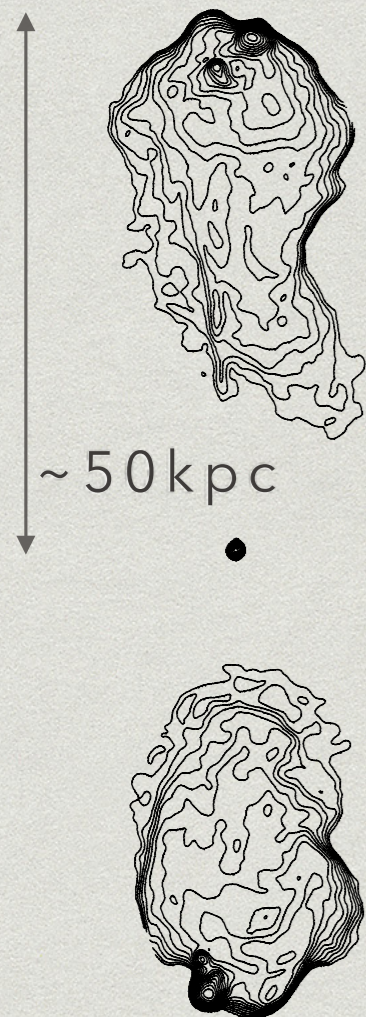
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Accretion rate...

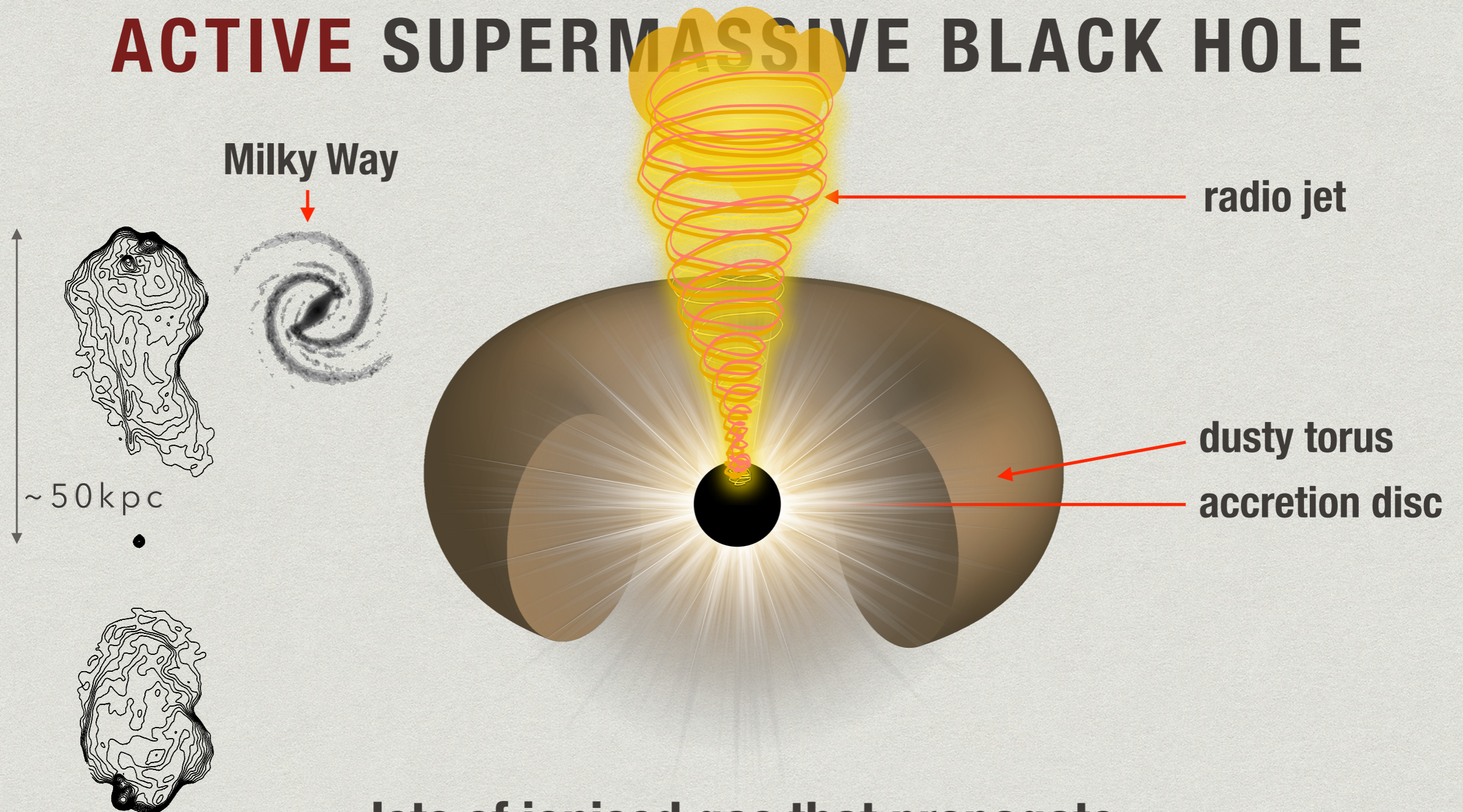
$$\dot{M} = \frac{L}{\eta c^2} = \frac{10^{10} L_{\odot}}{0.1 c^2} \approx 0.01 M_{\odot} \text{ yr}^{-1}$$

# ACTIVE SUPERMASSIVE BLACK HOLE



**Jets of ionised gas that propagate perpendicular to the plane of the accretion disc emit radio emission**

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# SUPERMASSIVE BLACK HOLE **FEEDBACK**

The gravitational energy of material falling into the centre of a galaxy toward the central supermassive black hole is released in the form of

- ▶ **radiation ranging from IR to X-rays**
- ▶ **mildly relativistic accretion winds**
- ▶ **relativistic radio jets**

Radiation, winds and jets “feed back” energy and momentum into the surrounding area of the galaxy

# WHY IS FEEDBACK **IMPORTANT**?

**Feedback may help suppress ongoing star formation**

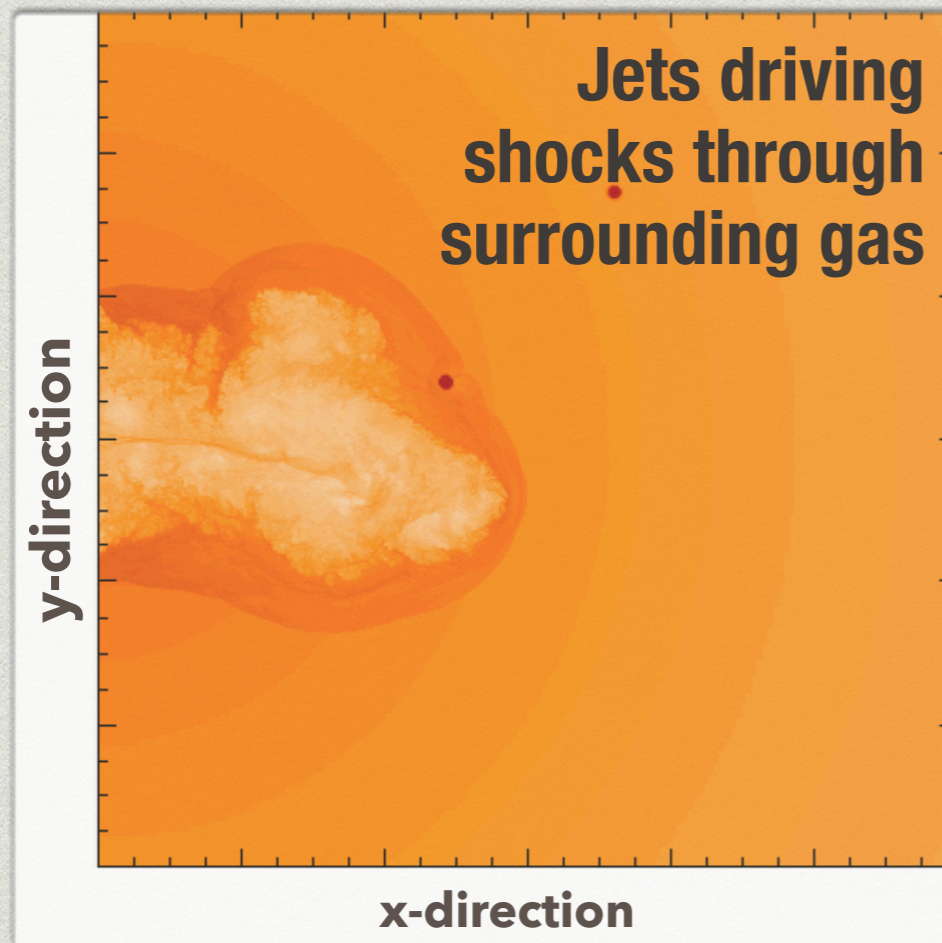


Image Credit: Vincenzo Antonuccio-Delogu

# WHY IS FEEDBACK IMPORTANT?

Feedback may help suppress ongoing star formation **and** reduce the overproduction of massive galaxies seen in models

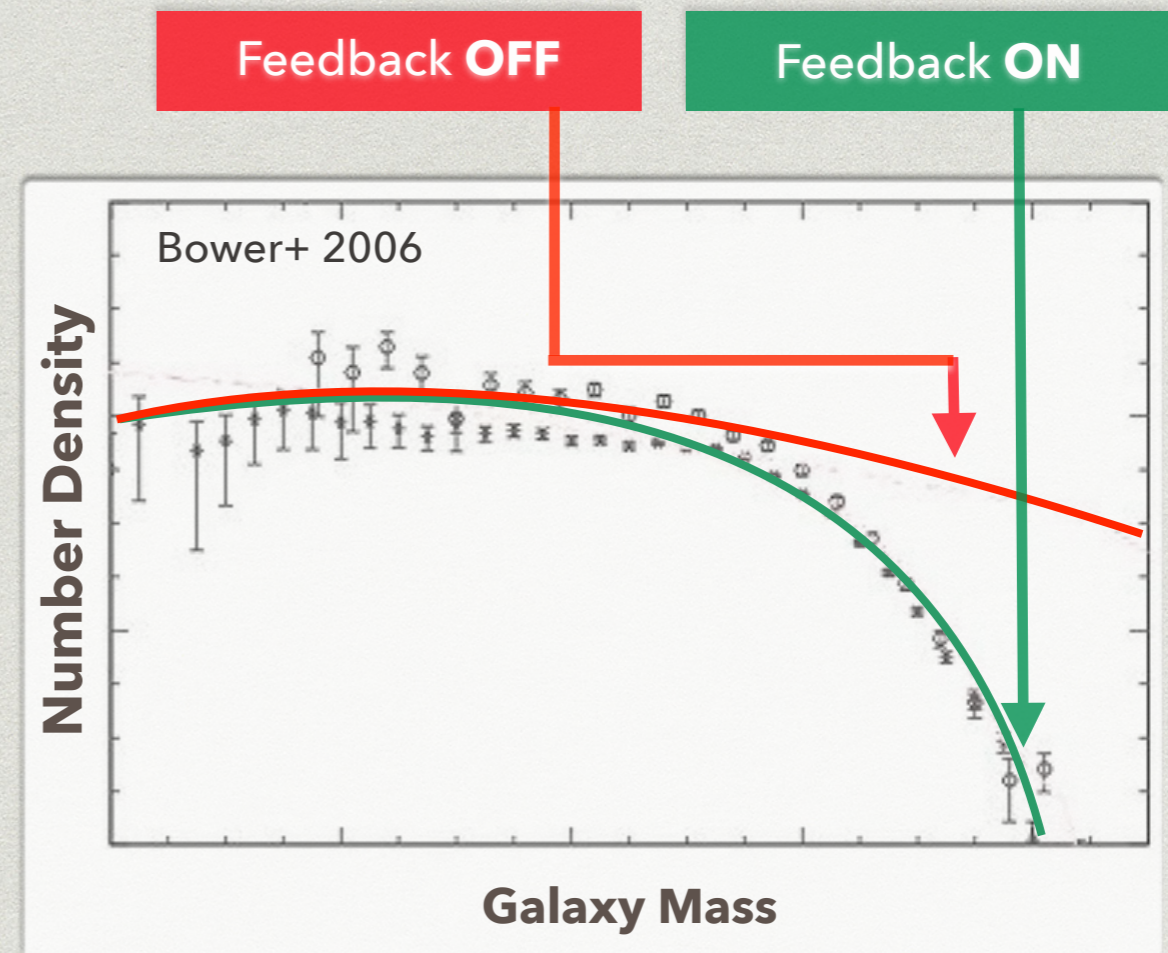
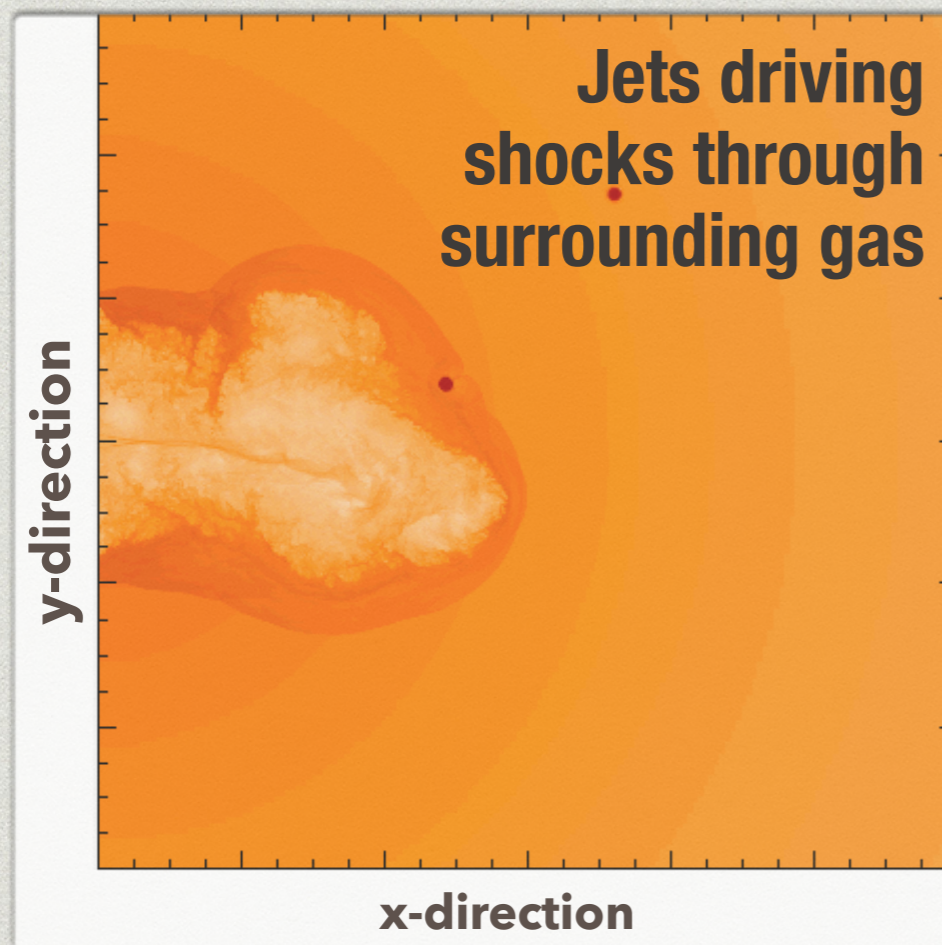


Image Credit: Vincenzo Antonuccio-Delogu

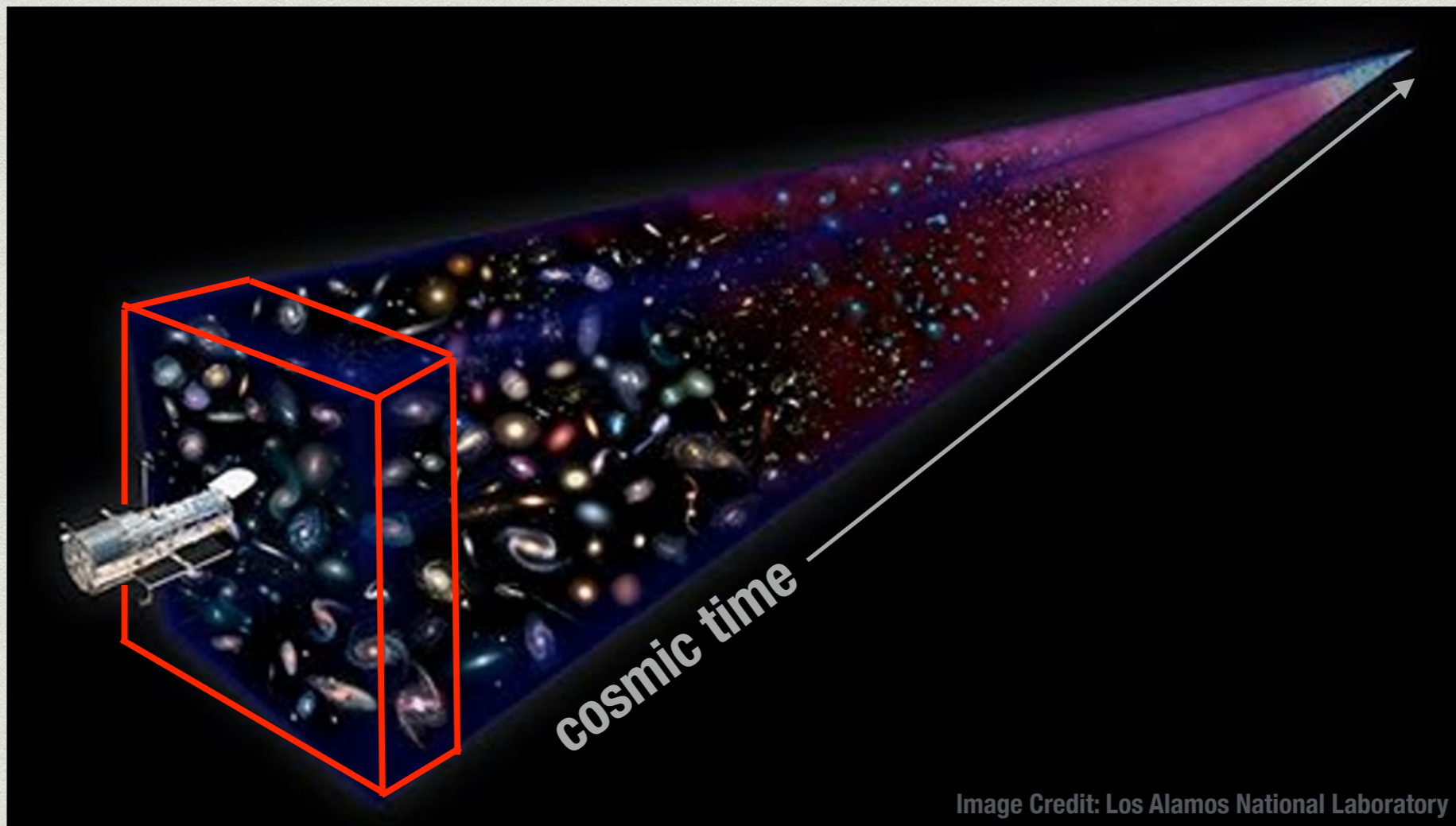
# **GOAL:**

**1) IDENTIFY A DIVERSE POPULATION OF GALAXIES HOSTING AN ACTIVE BLACK HOLE OVER A LARGE CHUNK OF COSMIC TIME**

**2) COMPARE THE STAR-FORMATION ACTIVITY OF THIS POPULATION WITH THOSE LACKING AN ACTIVE BLACK HOLE**

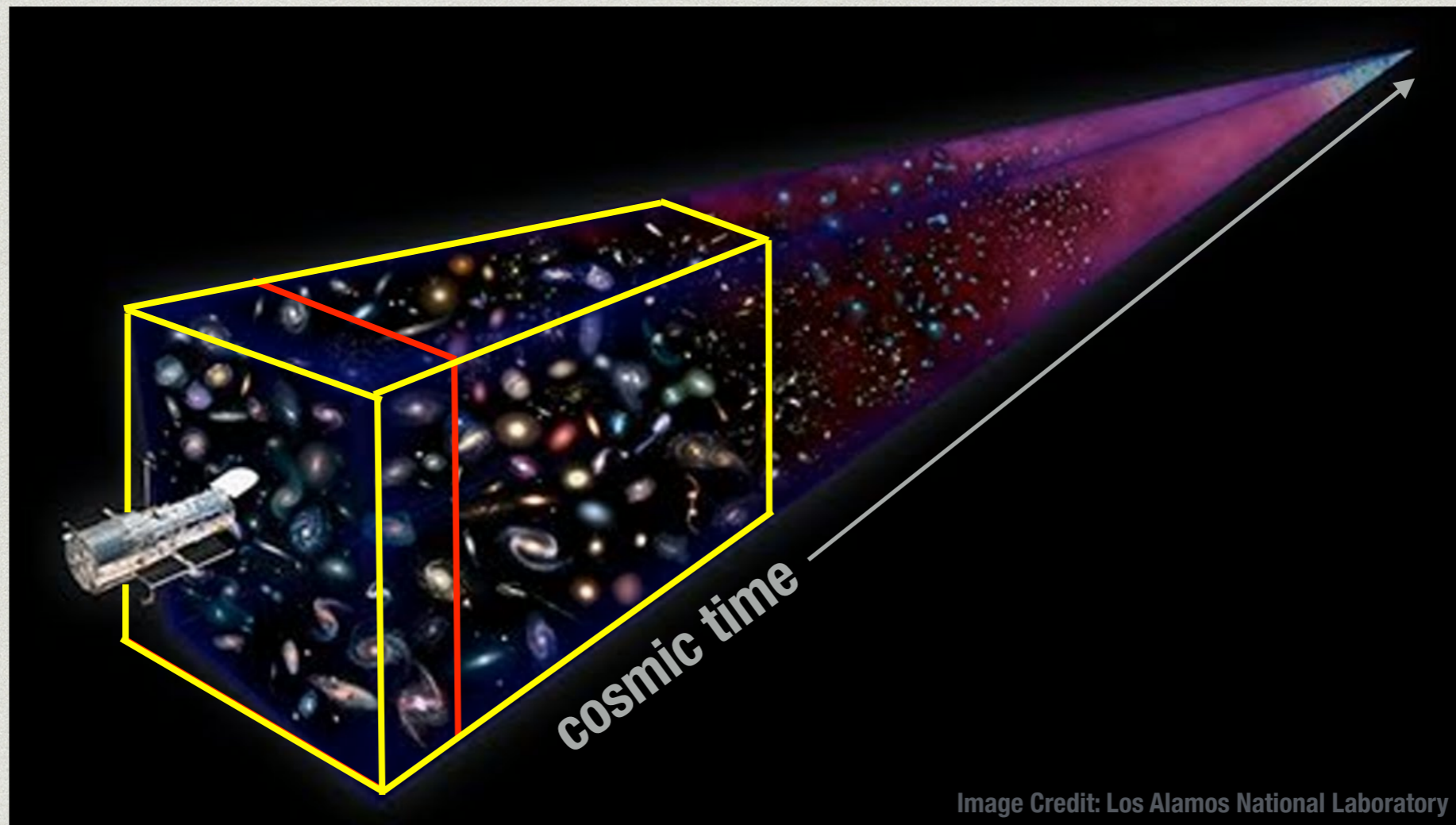
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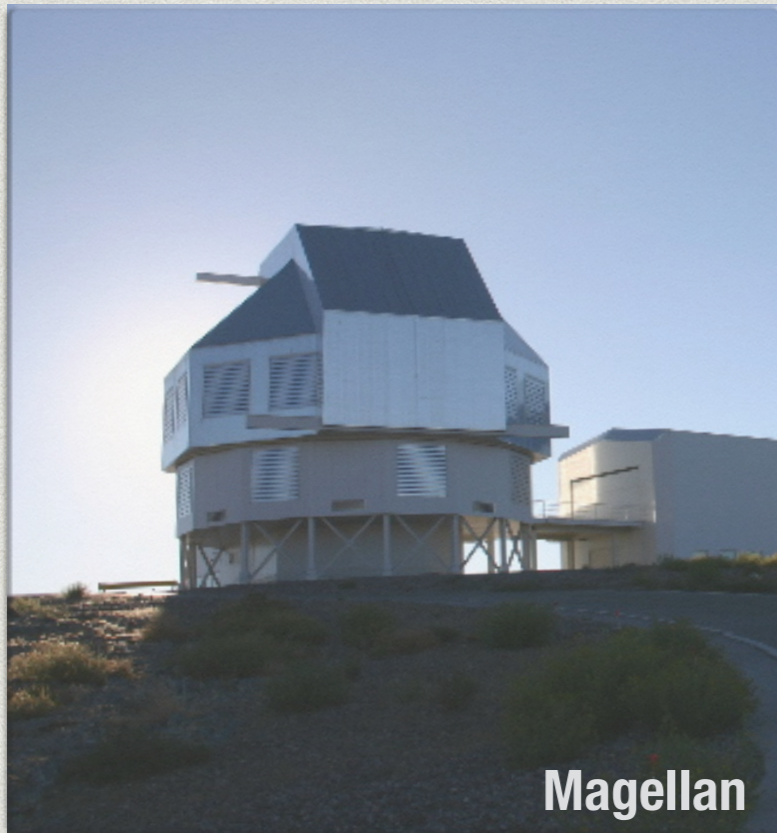


# GOAL:

1) IDENTIFY A POPULATION OF GALAXIES HOSTING AN ACTIVE BLACK HOLE OVER A LARGE CHUNK OF COSMIC TIME

## A multi-wavelength approach

Optical



Radio



Optical/IR



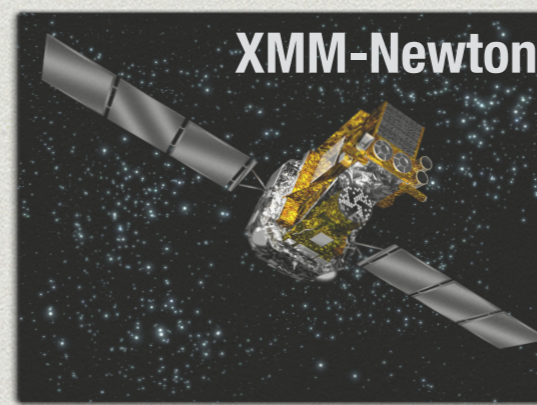
Infrared



Infrared



X-Ray



X-Ray



Infrared

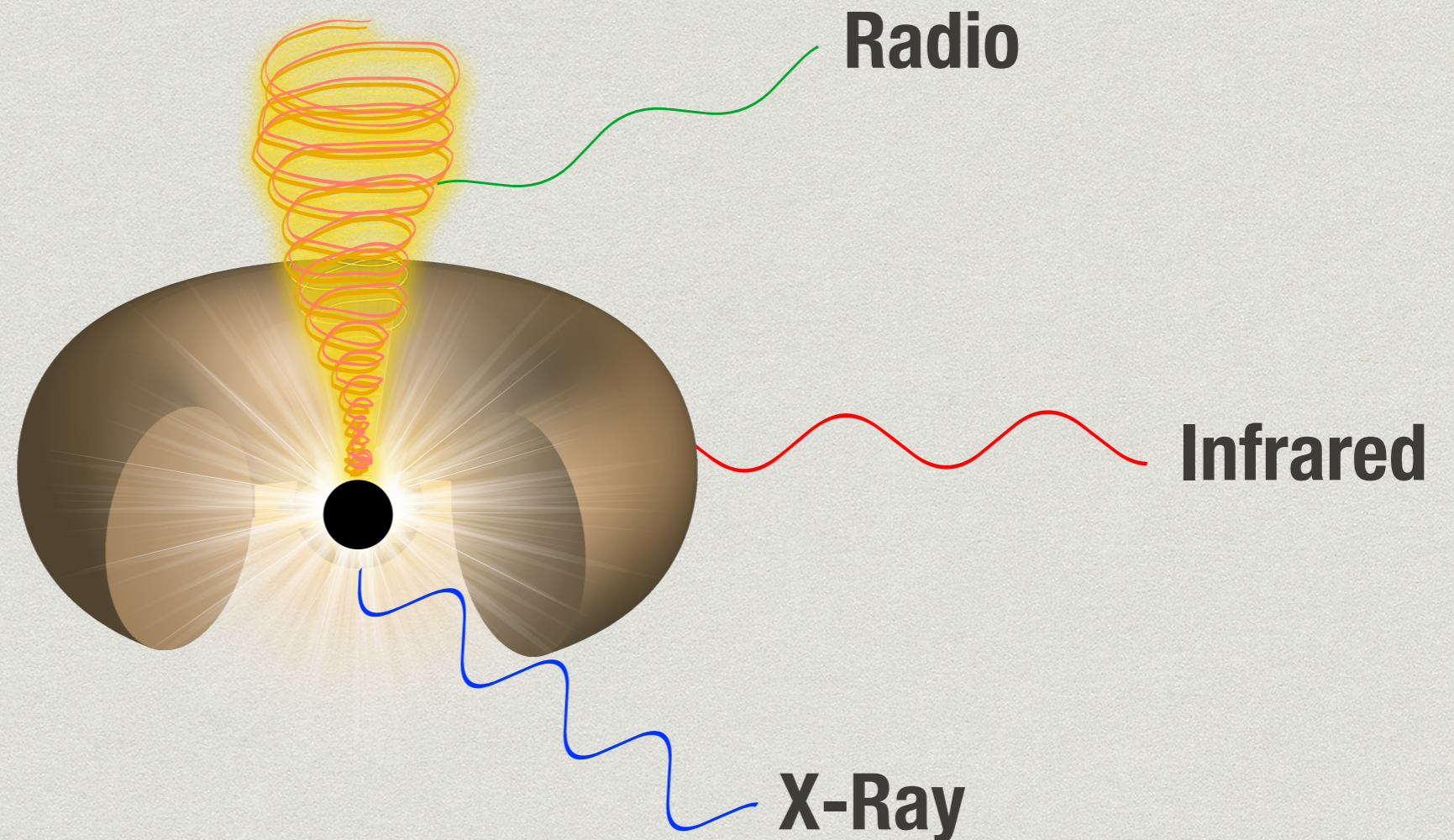
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# GOAL:

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Requires a  
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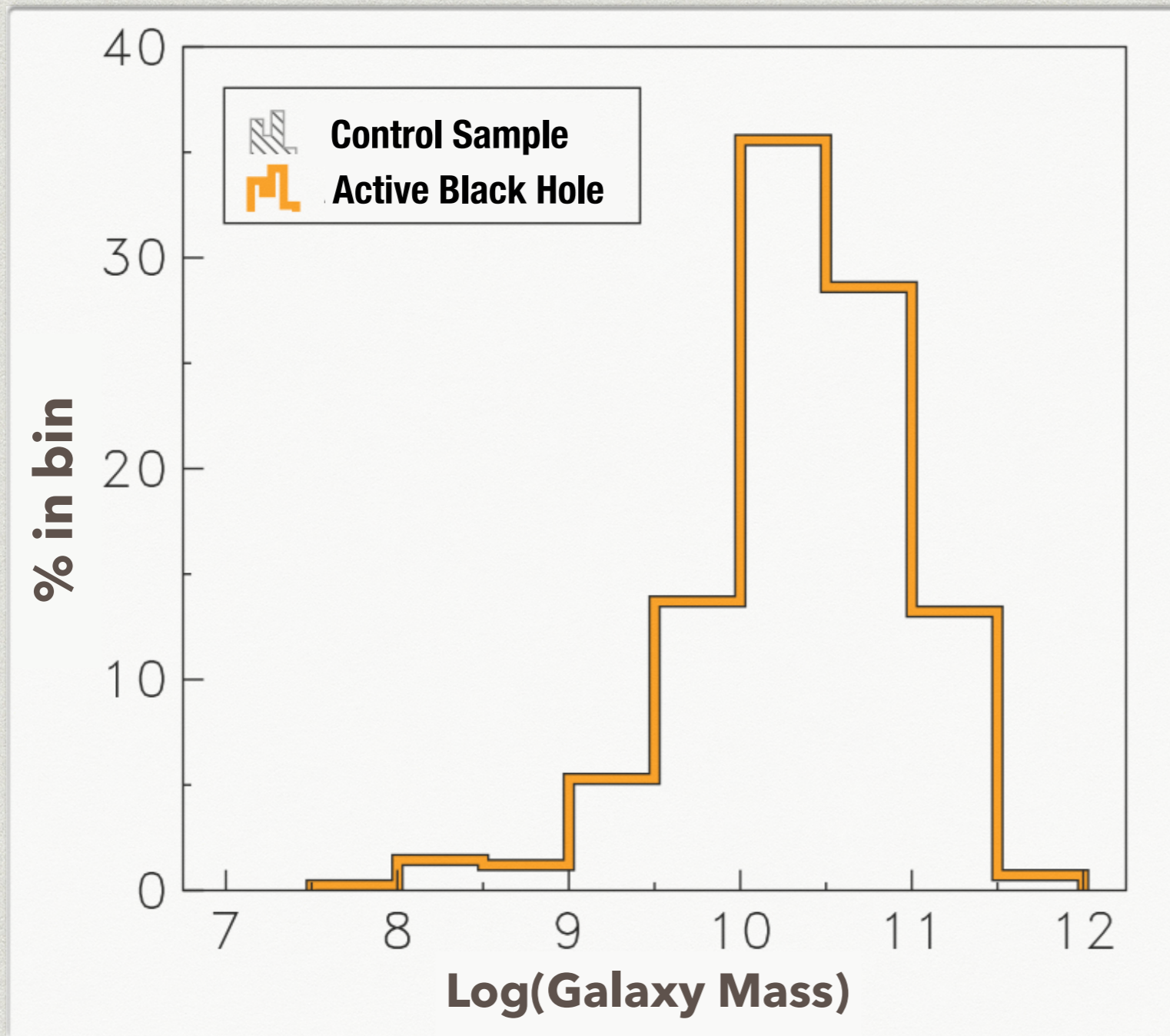


Control Sample

Image Credit: NASA/JPL-Caltech

# GOAL:

## 2) COMPARE THE STAR-FORMATION ACTIVITY OF GALAXIES HOSTING AN ACTIVE BLACK HOLE TO THOSE WITHOUT

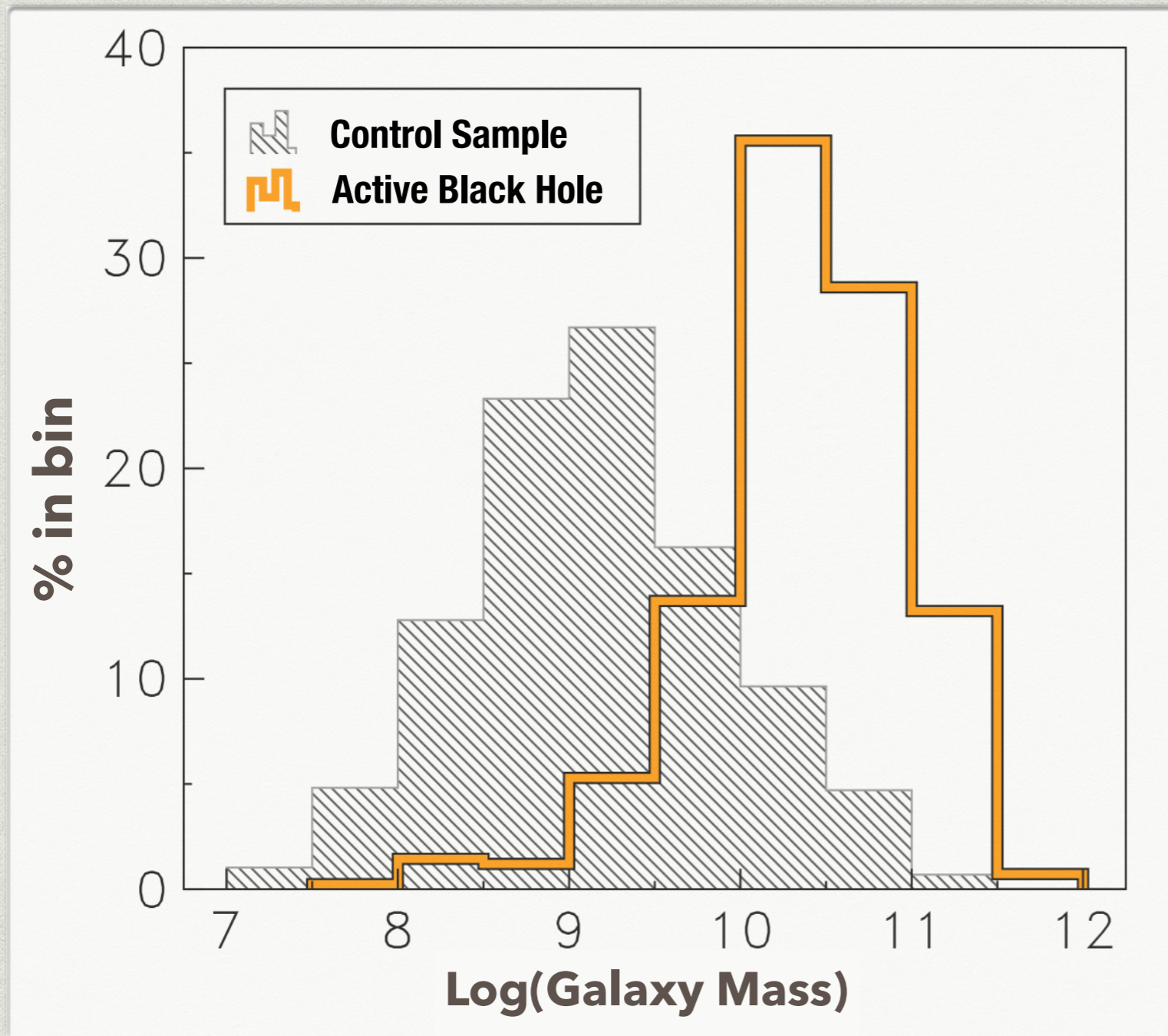


**Active supermassive black holes are preferentially hosted in galaxies with a high stellar mass (e.g., Aird+12)**

**A galaxy's stellar mass is tightly correlated with its star-formation rate (e.g., Noeske+07)**

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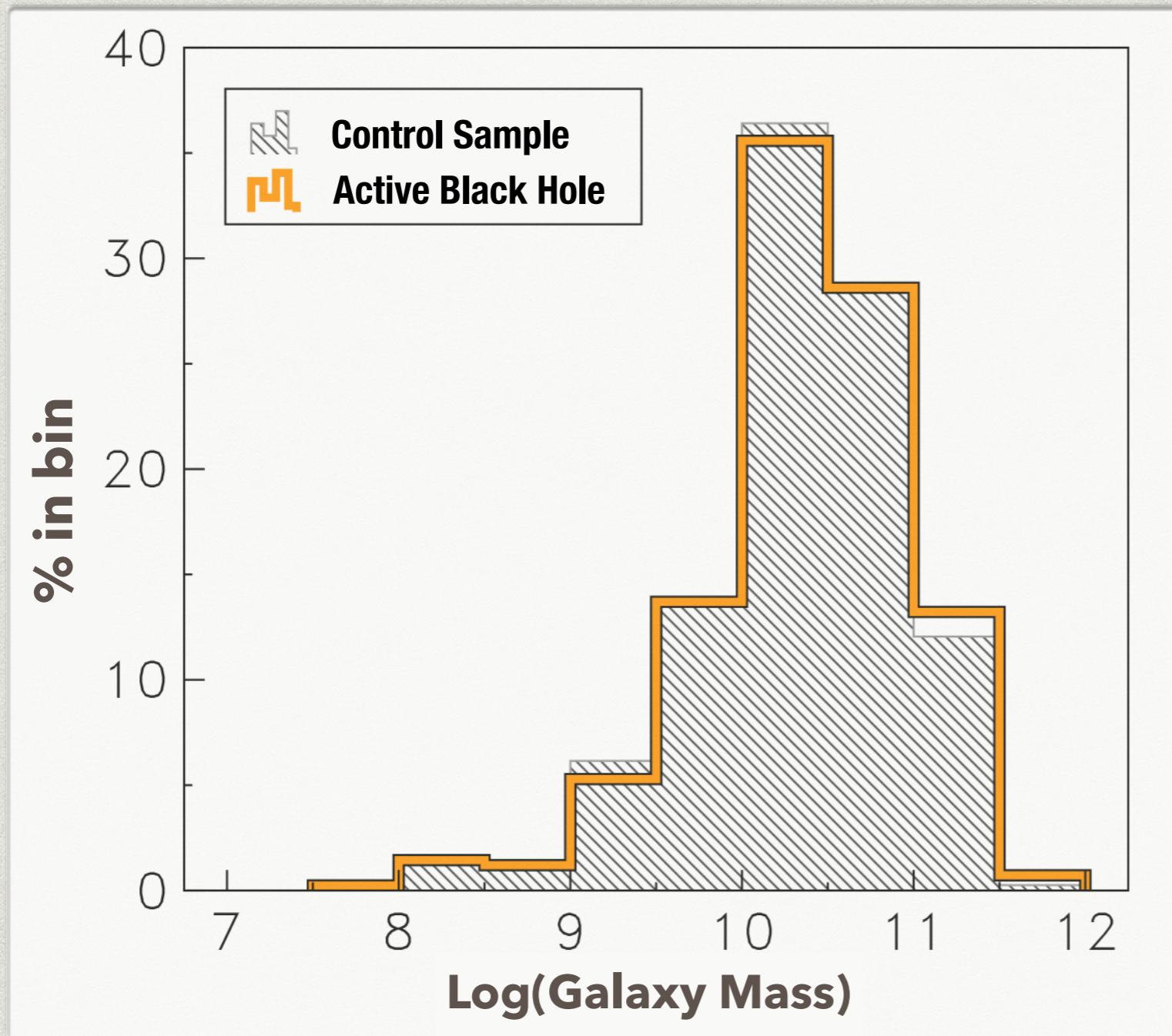


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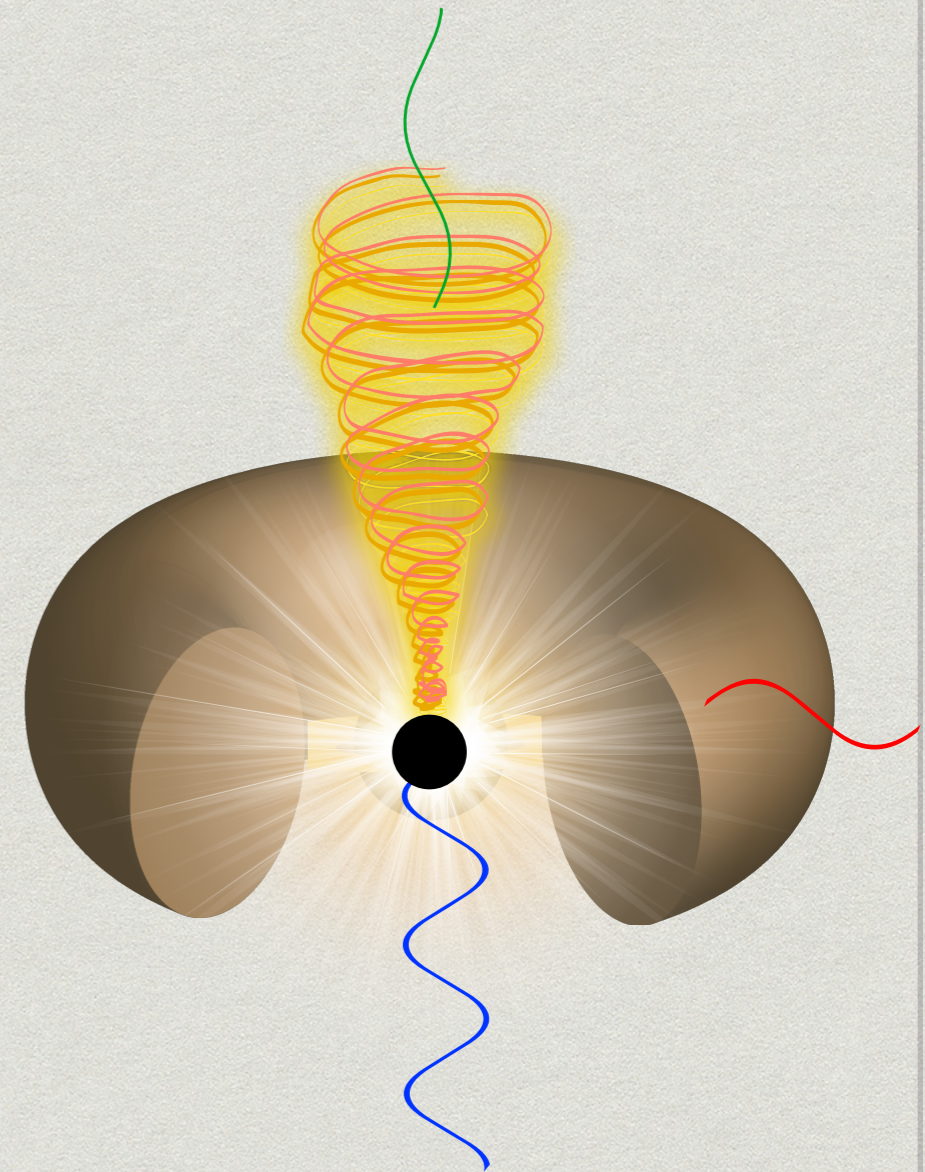
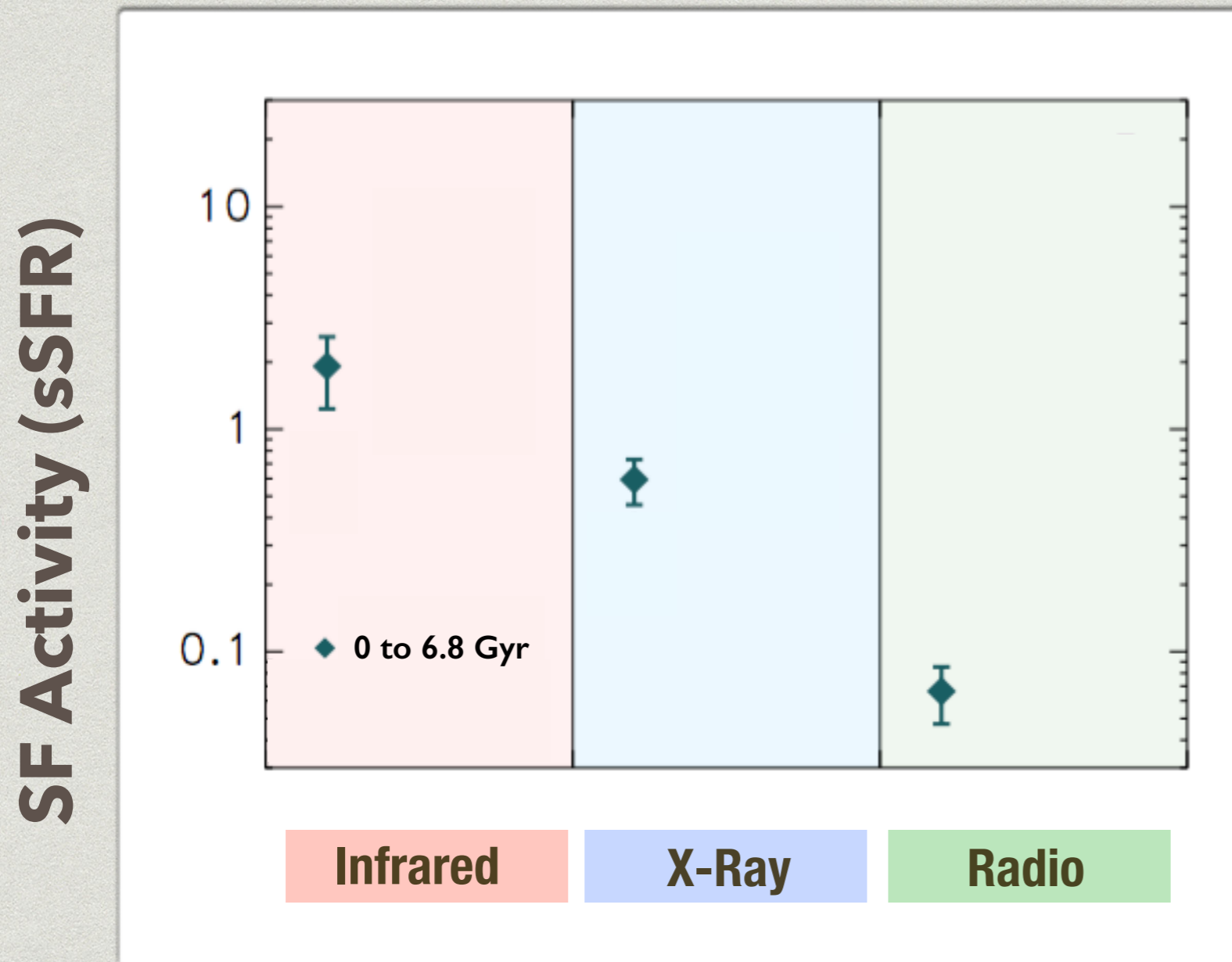


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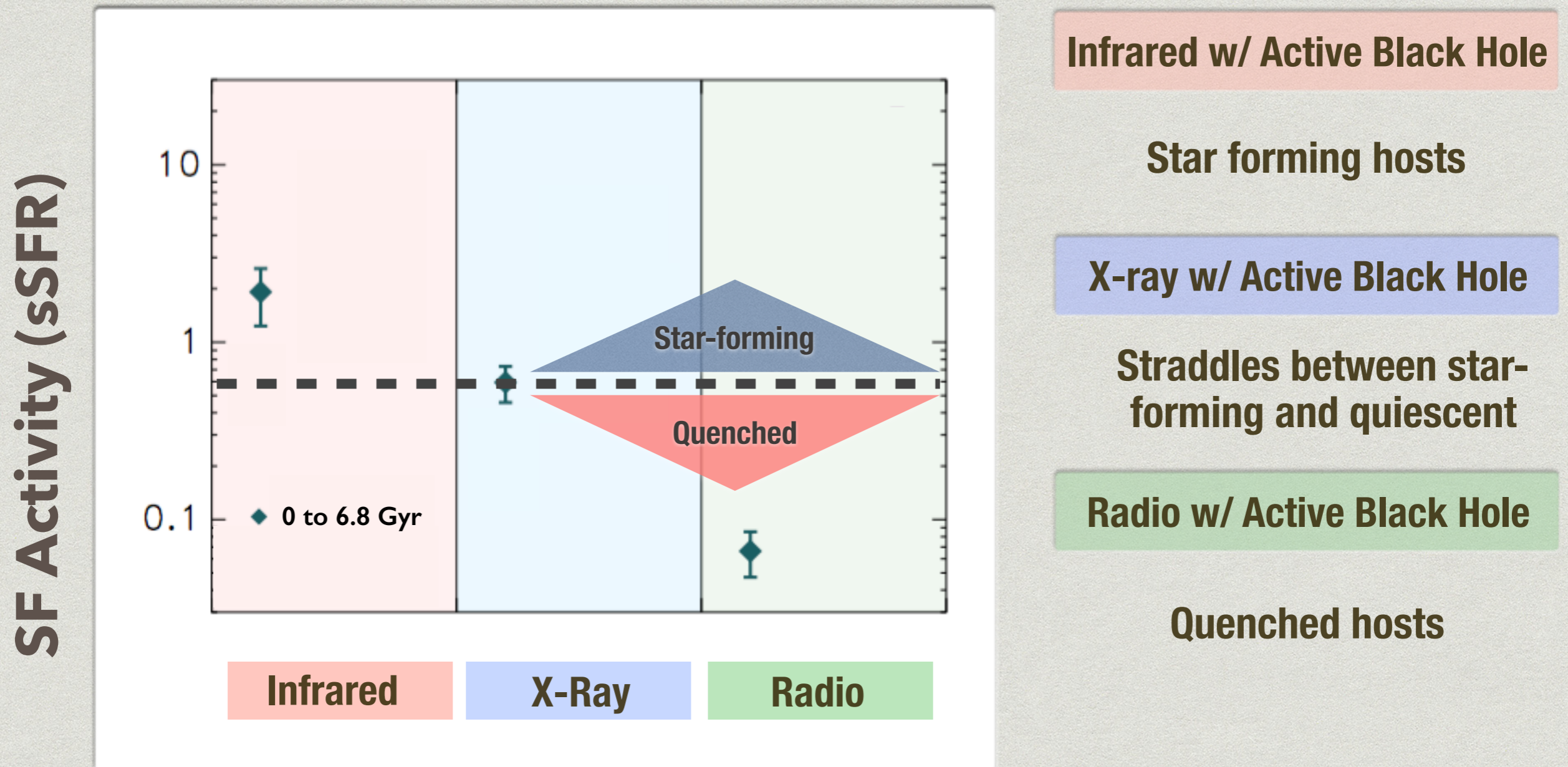
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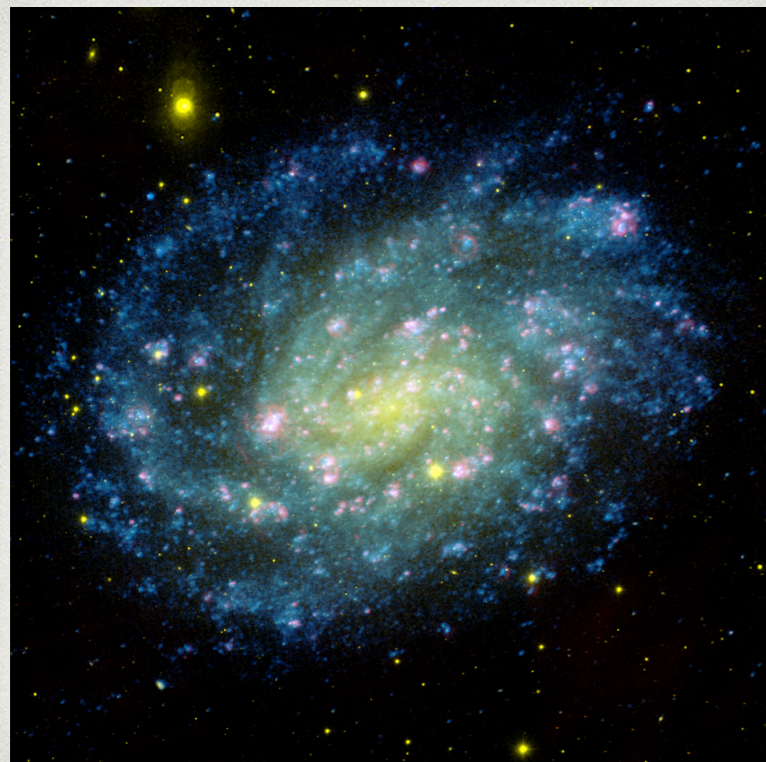
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# **GOAL:**

## **COMPARE THE STAR-FORMATION ACTIVITY OF GALAXIES HOSTING AN ACTIVE BLACK HOLE TO THOSE WITHOUT**

**Infrared w/ Active Black Hole**



**Younger Galaxies**

**X-Ray w/ Active Black Hole**



**Transitional**

**Radio w/ Active Black Hole**

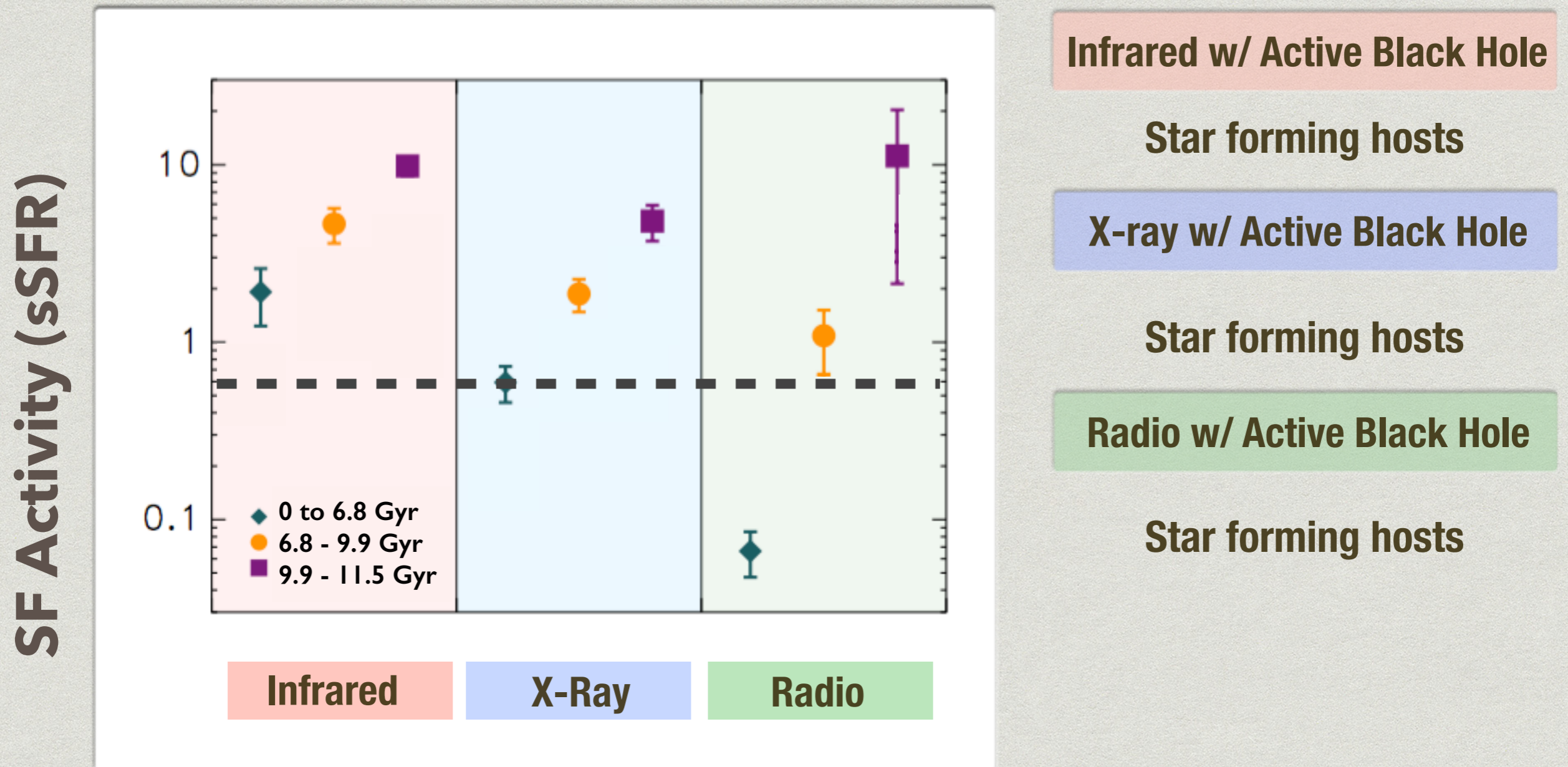


**Older Galaxies**

Image Credit: NASA/JPL-Caltech

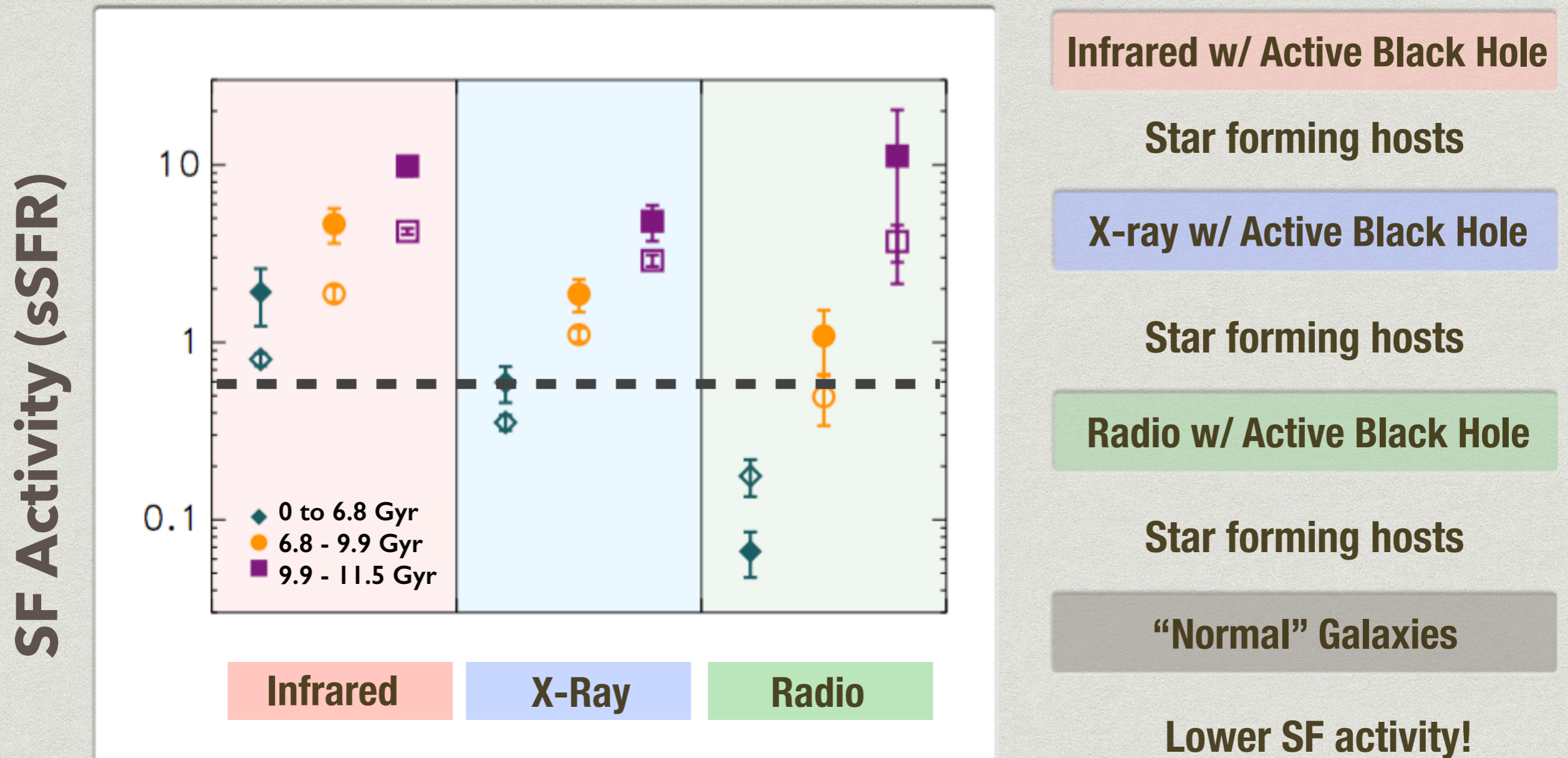
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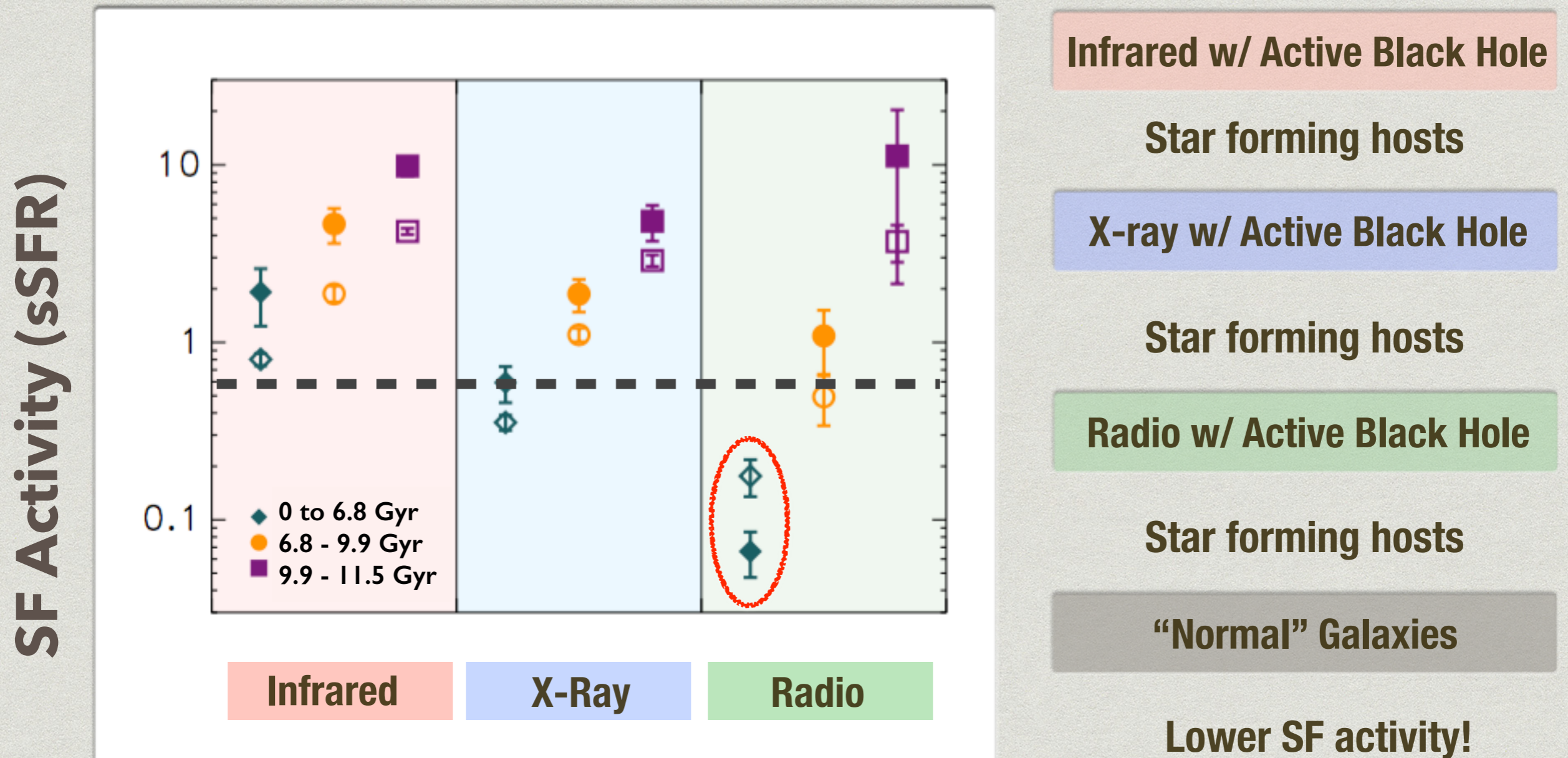
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# SUMMARY

- **The mean star-formation activity of the control sample or “normal” galaxies tends to be lower than galaxies hosting an active supermassive black hole**
  - **This calls into question the significance of black hole feedback being a primary mechanism for the suppression of star-formation in galaxies**
  - **The only instance the above is not true is the radio-selected galaxies in the “local” universe**
- ... what about the rest of cosmic time?**