

# DO SUPERMASSIVE BLACK HOLES IMPACT A GALAXY'S ABILITY TO FORM NEW STARS?

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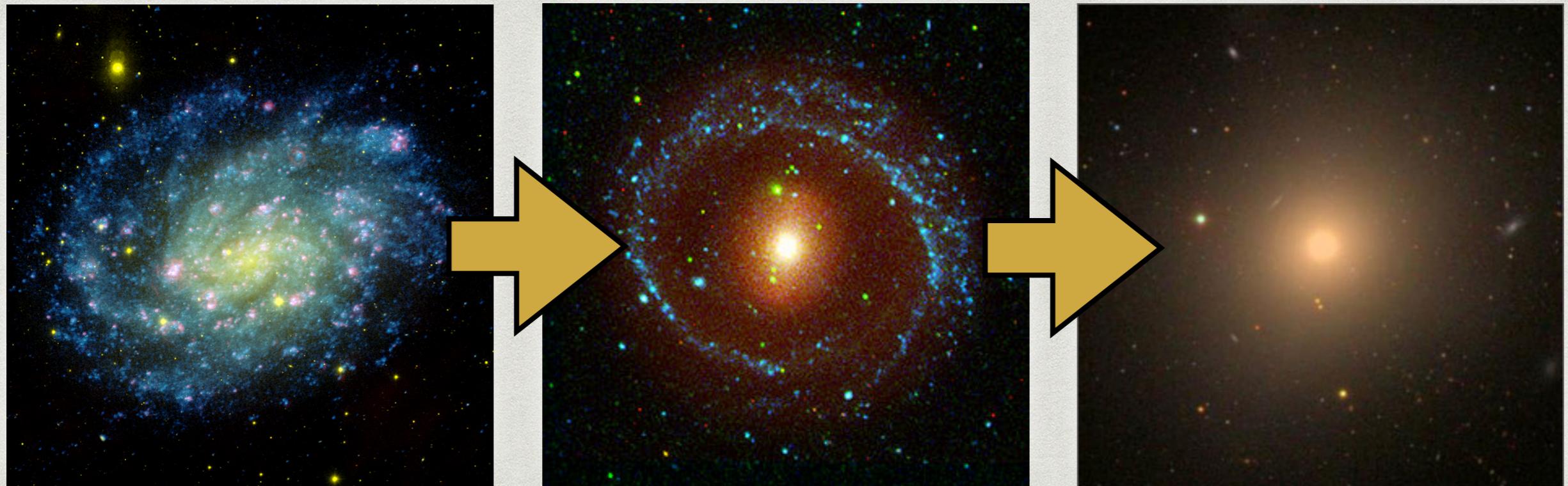
**MACQUARIE**  
University  
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Department of Industry and Science



# WHAT SUPPRESSES STAR-FORMATION IN GALAXIES?



Younger Galaxies

Transitional

Older Galaxies

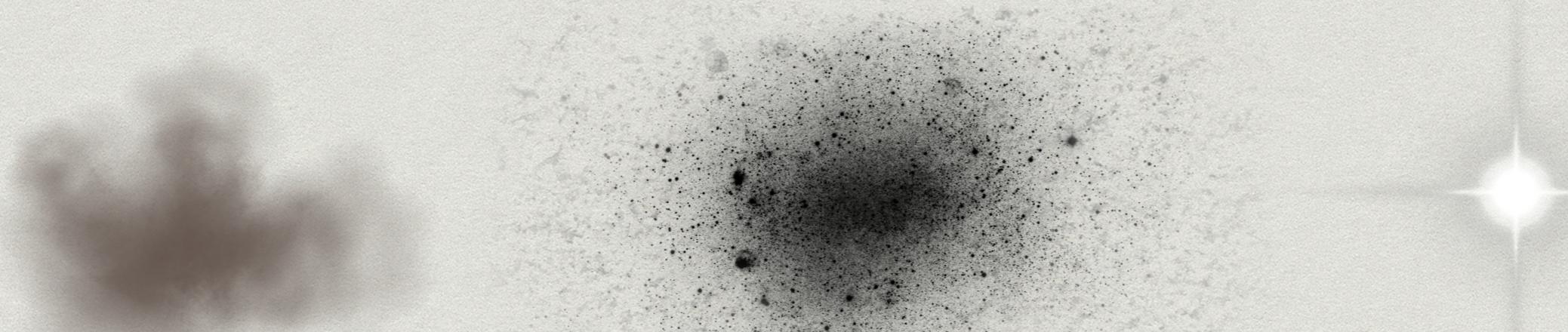
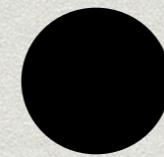
Understanding this is central to building a  
coherent picture of galaxy evolution

Image Credit: NASA/JPL-Caltech

# SUPERMASSIVE BLACK HOLE



# SUPERMASSIVE BLACK HOLE



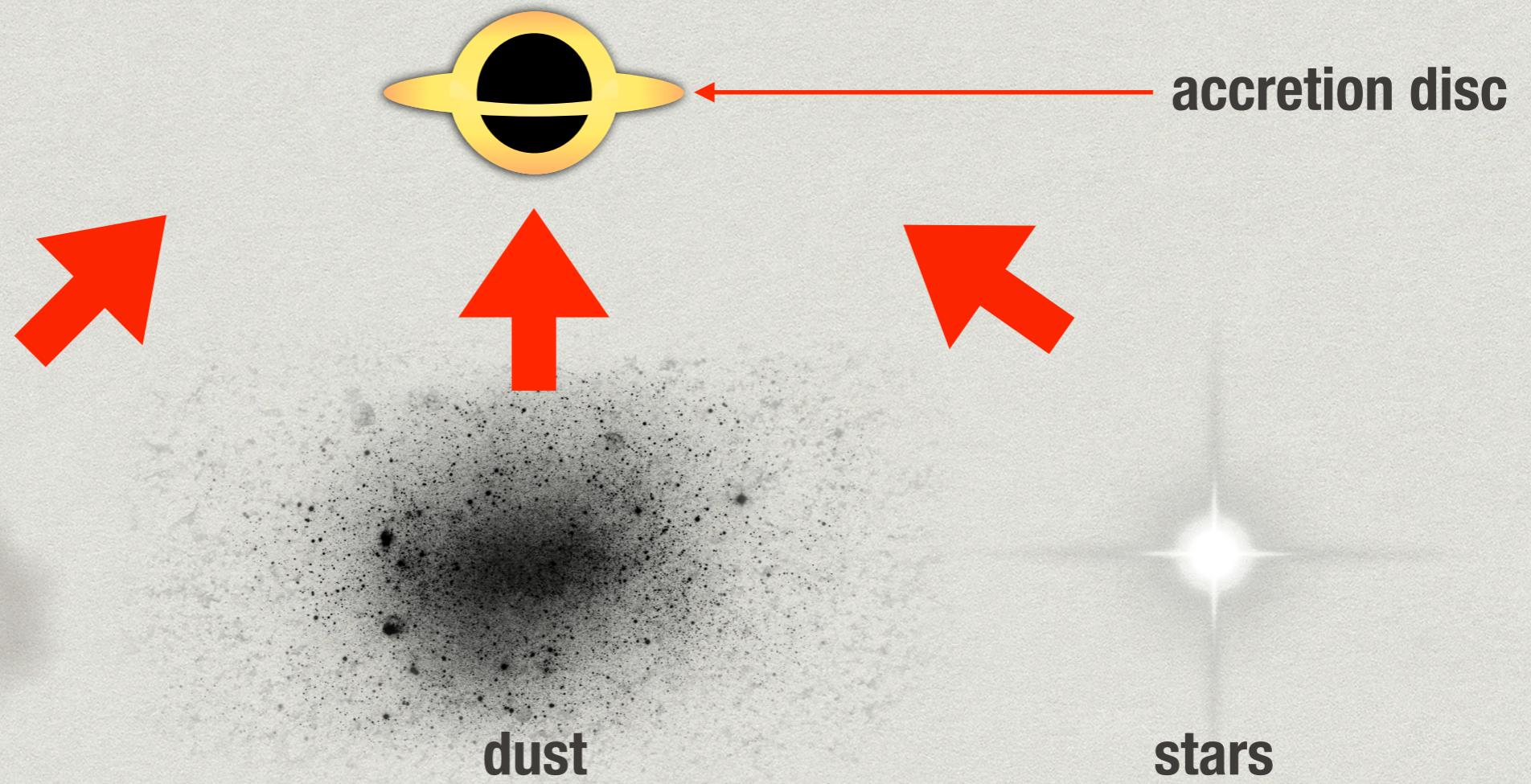
gas

dust

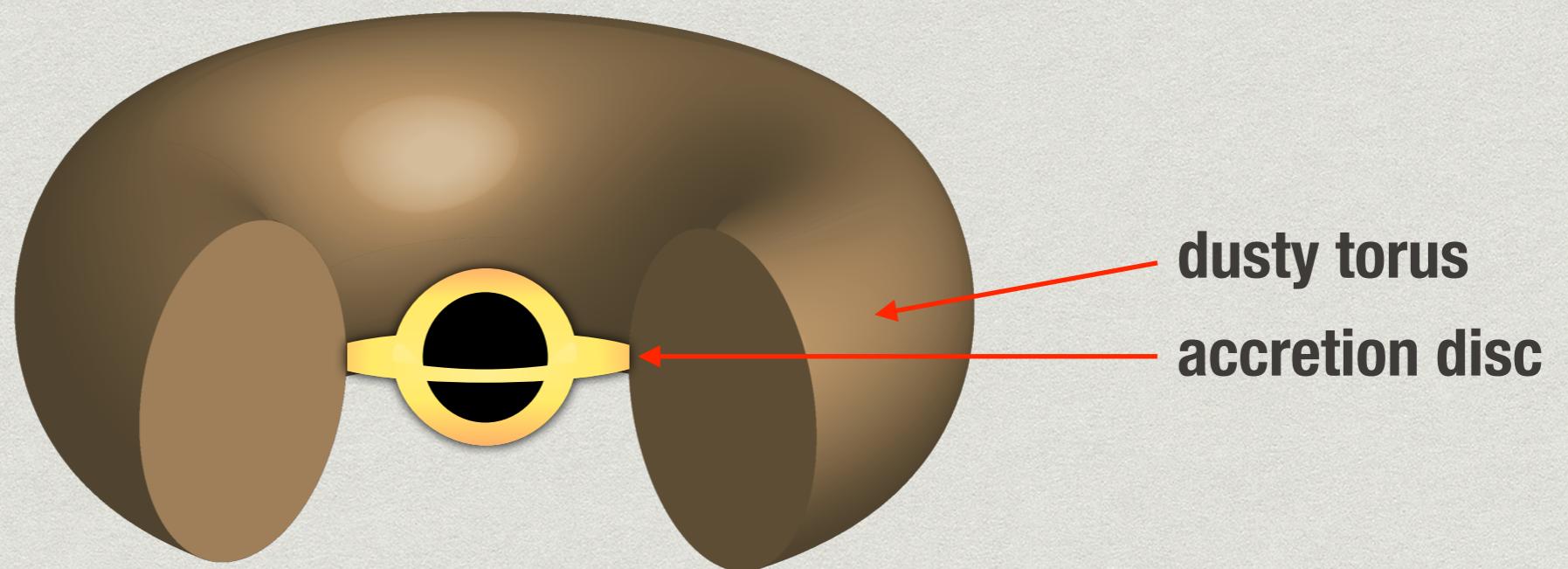
stars

# FEEDING A SUPERMASSIVE BLACK HOLE

Conservation of angular momentum causes fuel to form a disk as it spirals in

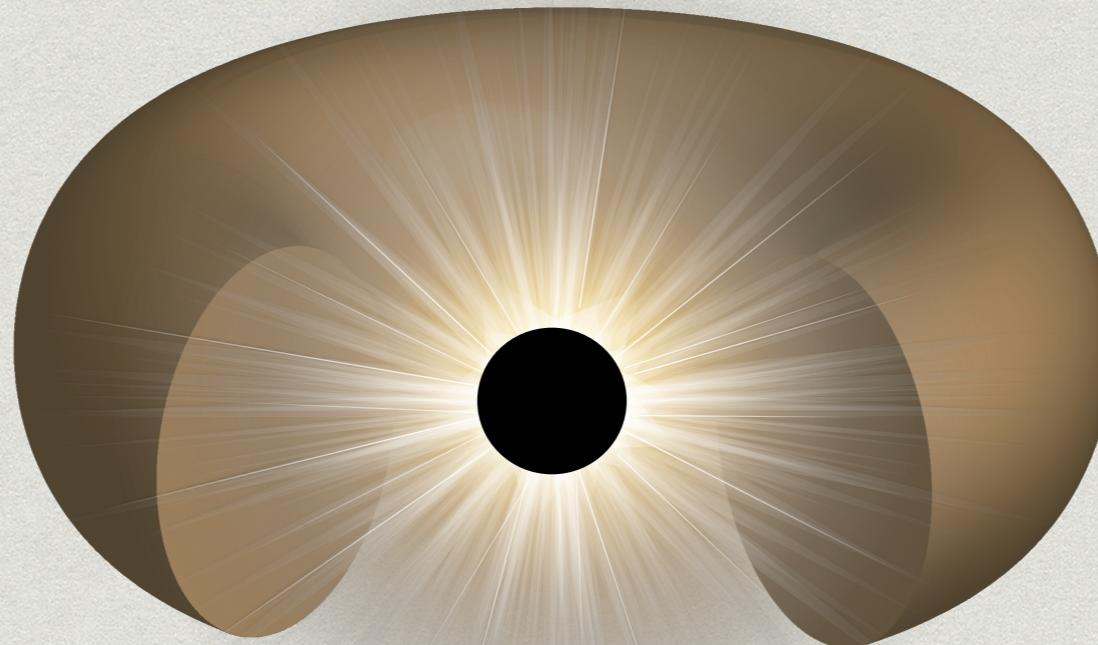


# FEEDING A SUPERMASSIVE BLACK HOLE



# ACTIVE SUPERMASSIVE BLACK HOLE

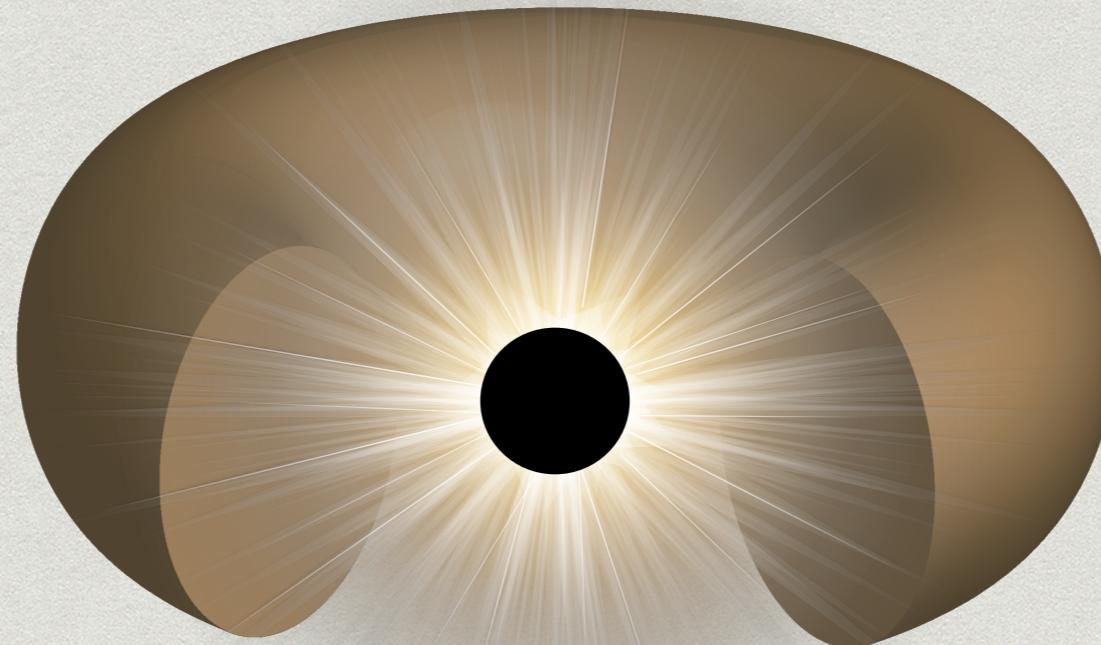
$L \sim 10^{37-41} \text{W}$   
(i.e.  $\sim 10^{10-14} L_\odot$ )



**Within this disc, there are dissipative processes, e.g. collisions, shocks, etc. This dissipated energy emerges as radiation**

# ACTIVE SUPERMASSIVE BLACK HOLE

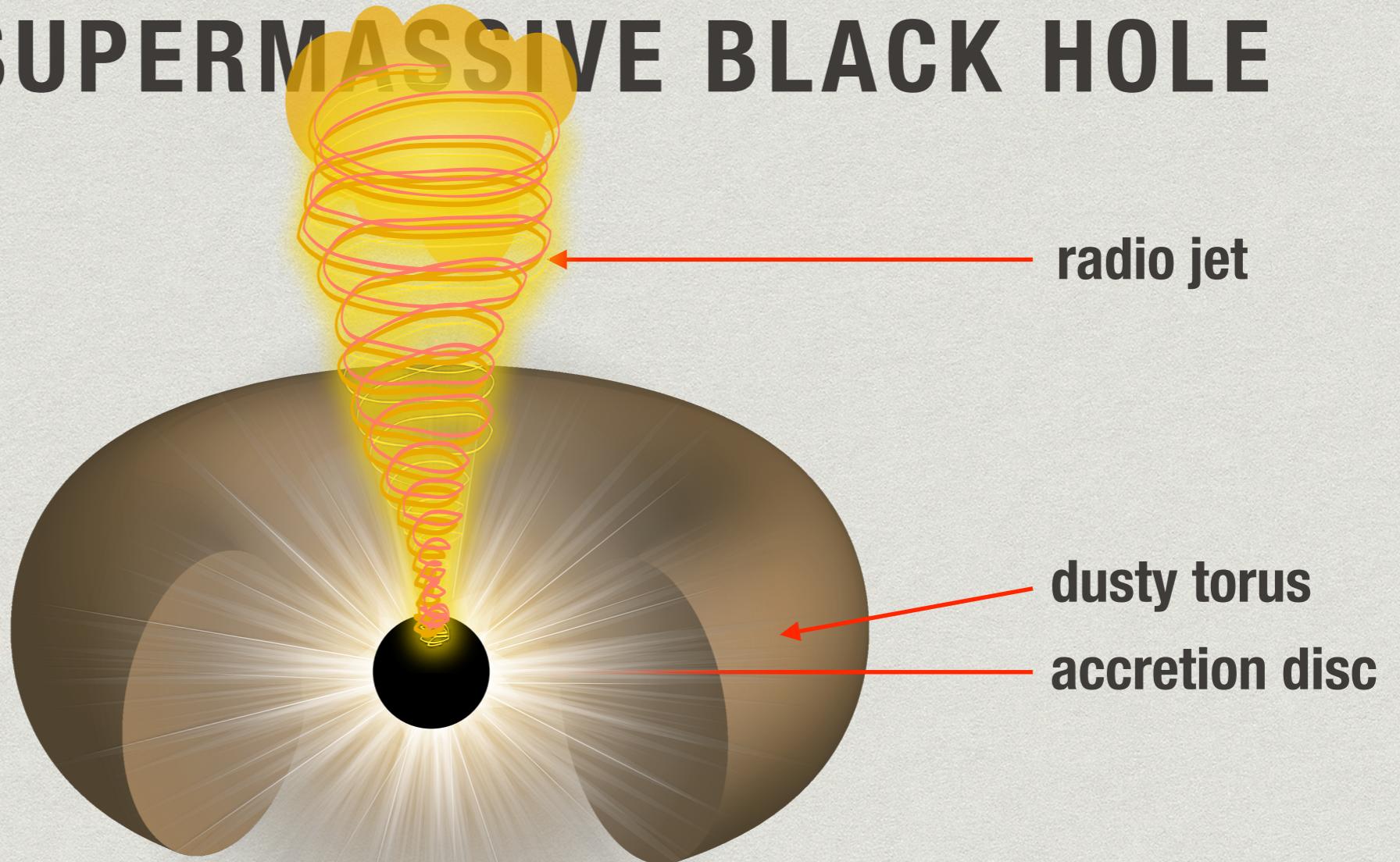
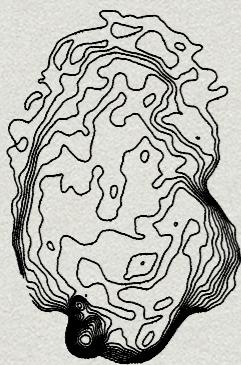
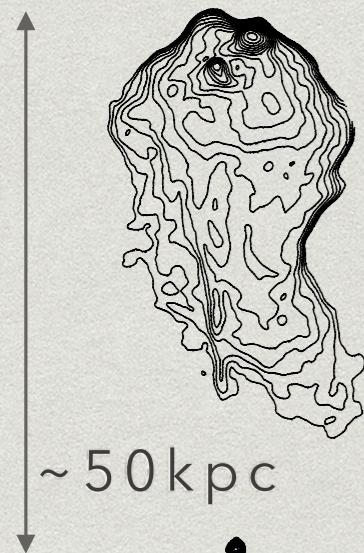
$L \sim 10^{37-41} \text{W}$   
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Accretion rate...

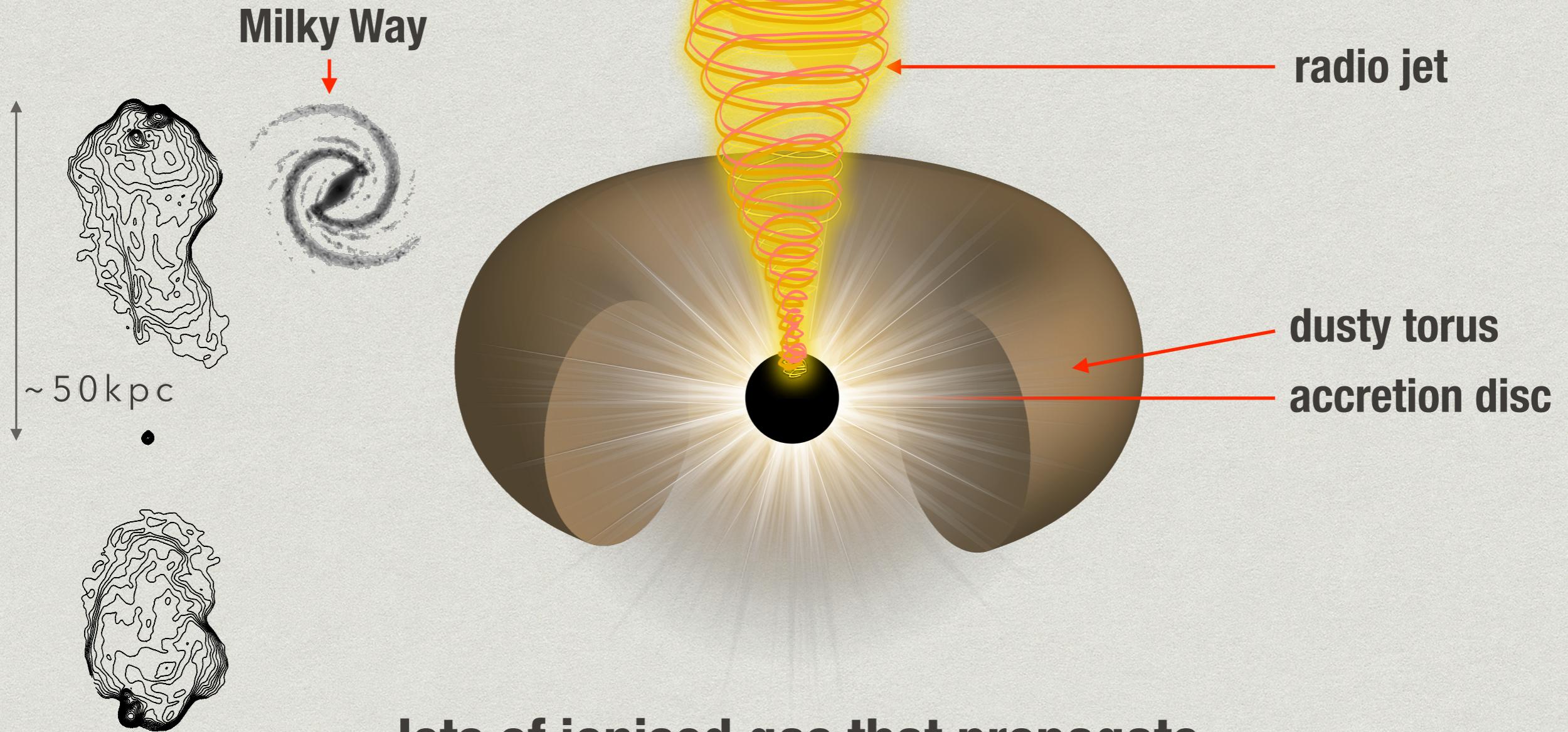
$$\dot{M} = \frac{L}{\eta c^2} = \frac{10^{10} L_\odot}{0.1 c^2} \approx 0.01 M_\odot \text{yr}^{-1}$$

# ACTIVE SUPERMASSIVE BLACK HOLE



**Jets of ionised gas that propagate perpendicular to the plane of the accretion disc emit radio emission**

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# SUPERMASSIVE BLACK HOLE FEEDBACK

The gravitational energy of material falling into the centre of a galaxy toward the central supermassive black hole is released in the form of

- ▶ **radiation ranging from IR to X-rays**
- ▶ **mildly relativistic accretion winds**
- ▶ **relativistic radio jets**

Radiation, winds and jets “feed back” energy and momentum into the surrounding area of the galaxy

# WHY IS FEEDBACK IMPORTANT?

Feedback may help suppress ongoing star formation

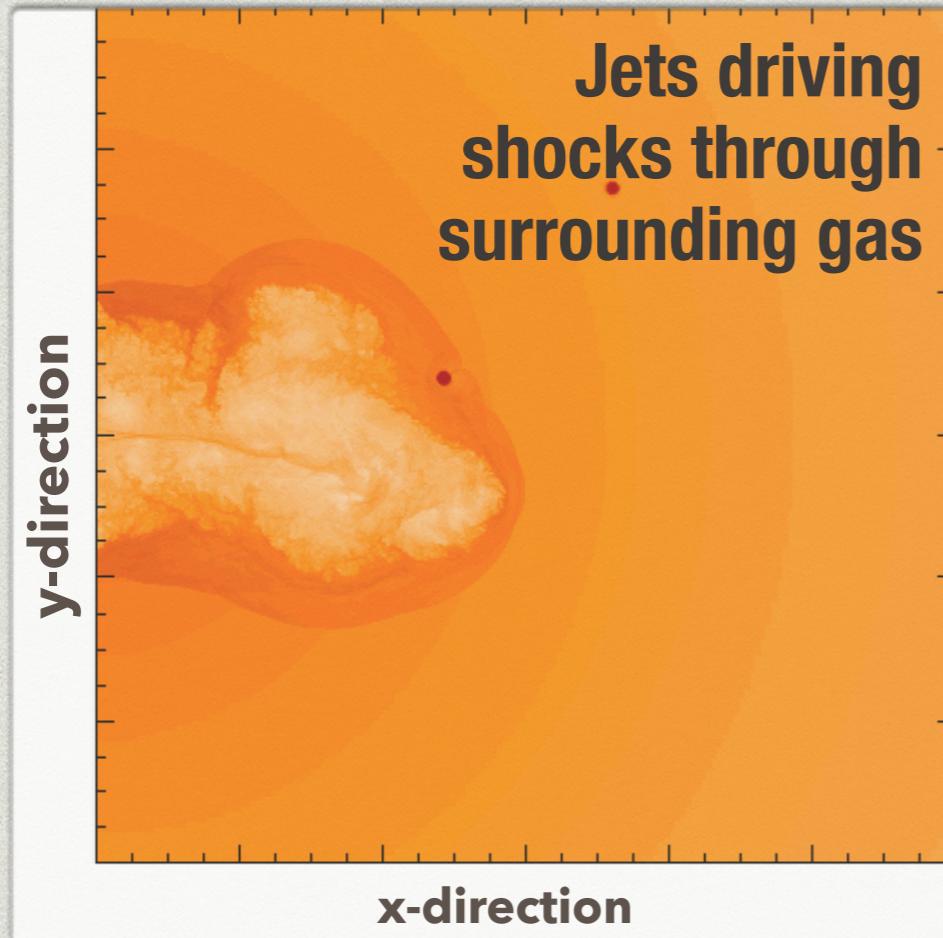


Image Credit: Vincenzo Antonuccio-Delogu

# WHY IS FEEDBACK IMPORTANT?

Feedback may help suppress ongoing star formation **and** reduce the overproduction of massive galaxies seen in models

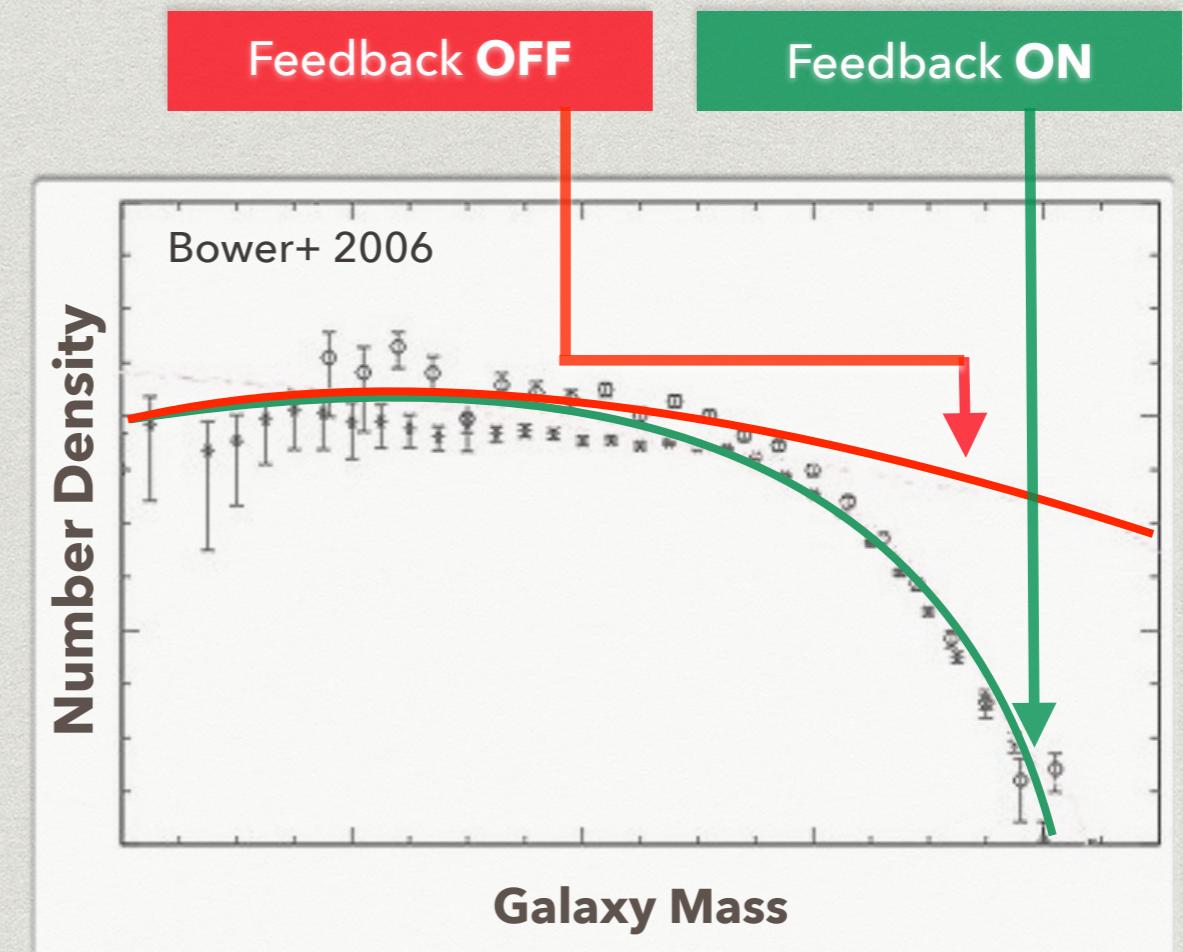
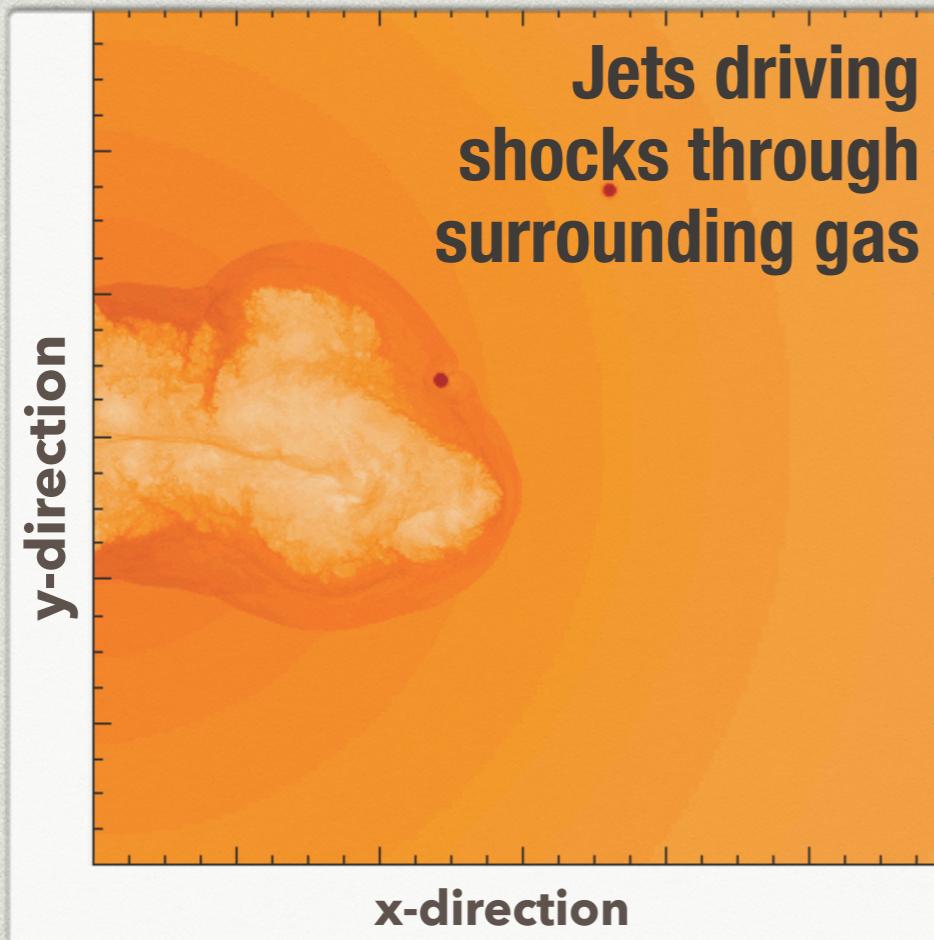


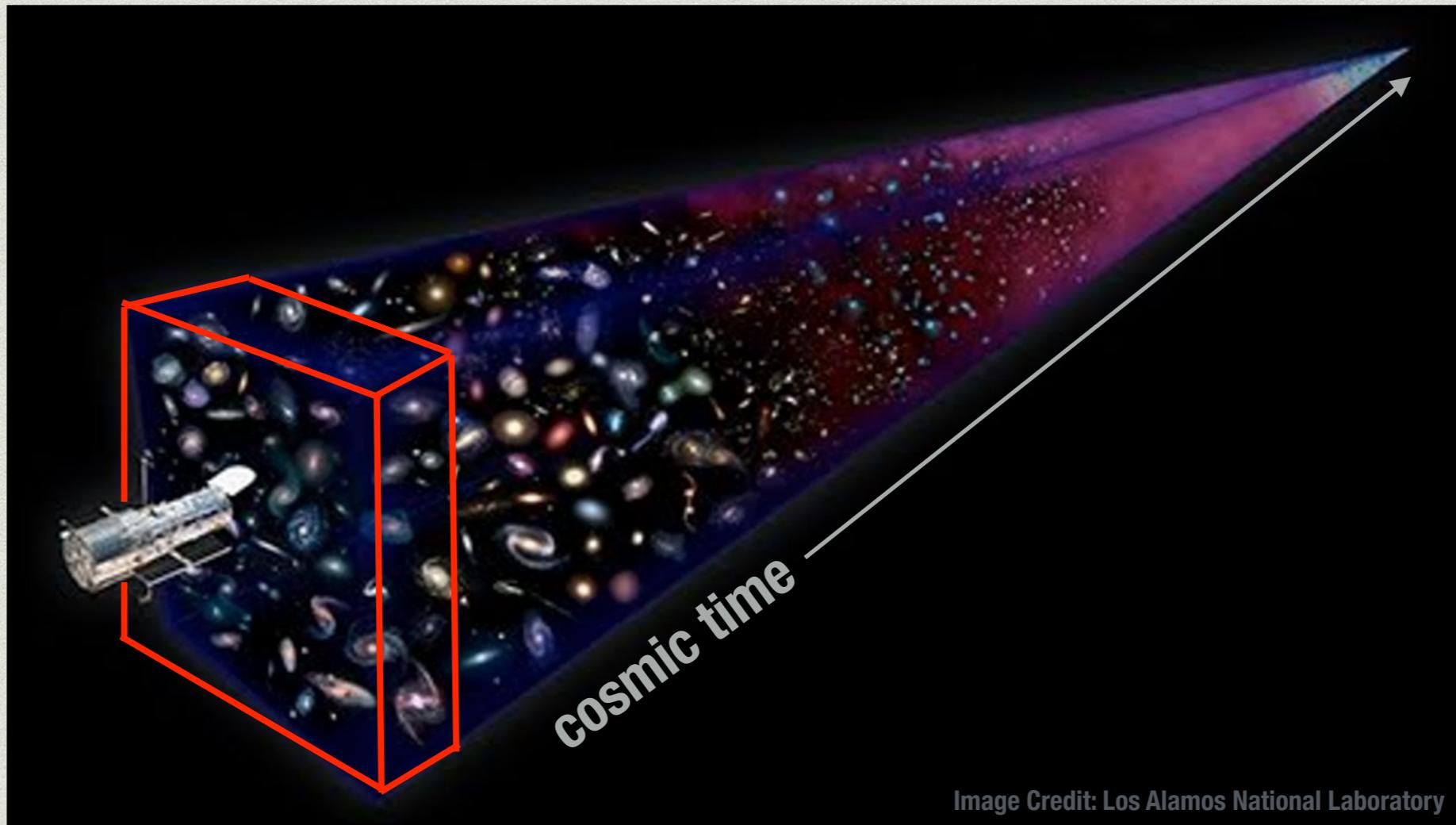
Image Credit: Vincenzo Antonuccio-Delogu

# GOAL:

- 1) IDENTIFY A DIVERSE POPULATION OF GALAXIES HOSTING AN ACTIVE BLACK HOLE OVER A LARGE CHUNK OF COSMIC TIME**
- 2) COMPARE THE STAR-FORMATION ACTIVITY OF THIS POPULATION WITH THOSE LACKING AN ACTIVE BLACK HOLE**

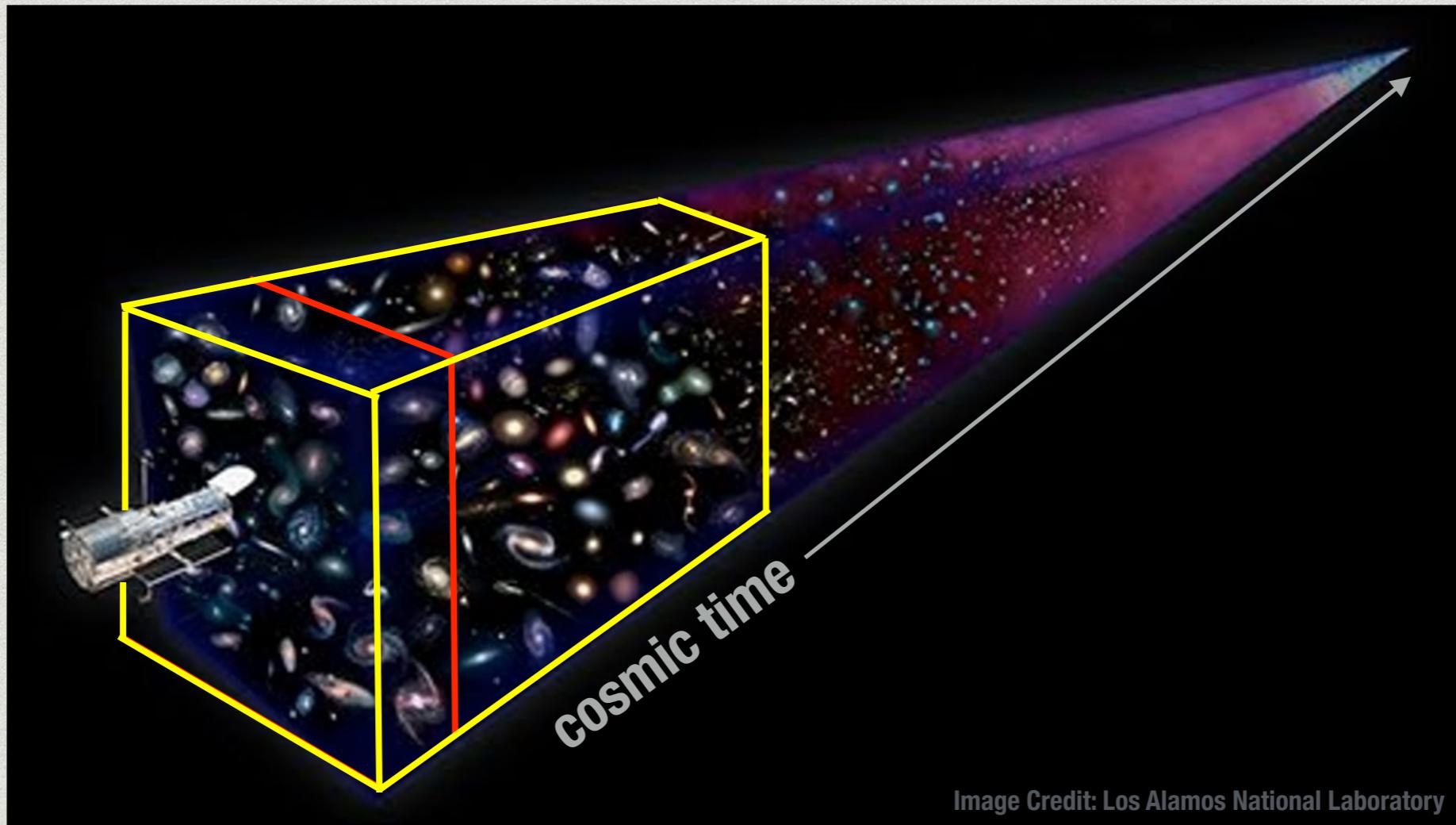
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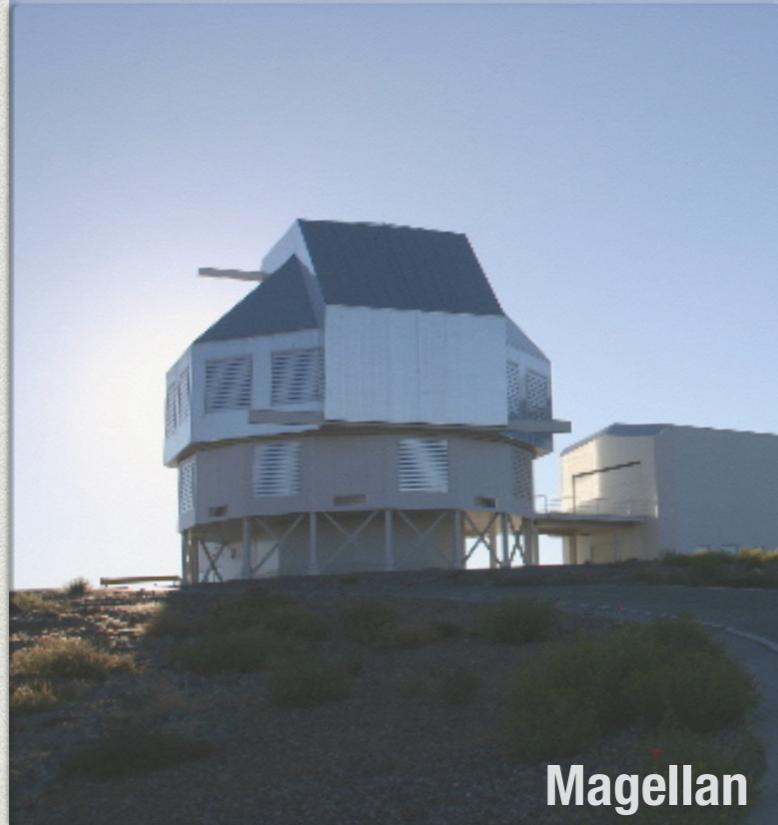


# GOAL:

1) IDENTIFY A POPULATION OF GALAXIES HOSTING AN ACTIVE BLACK HOLE OVER A LARGE CHUNK OF COSMIC TIME

A **multi**-wavelength approach

Optical



Magellan

Radio



VLA

Optical/IR



HST

Infrared



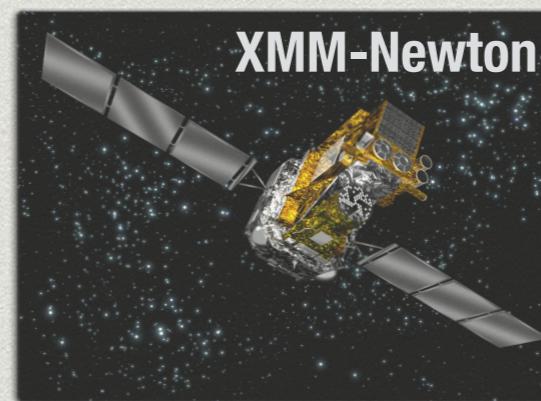
Spitzer

Infrared



Herschel

X-Ray



XMM-Newton

X-Ray

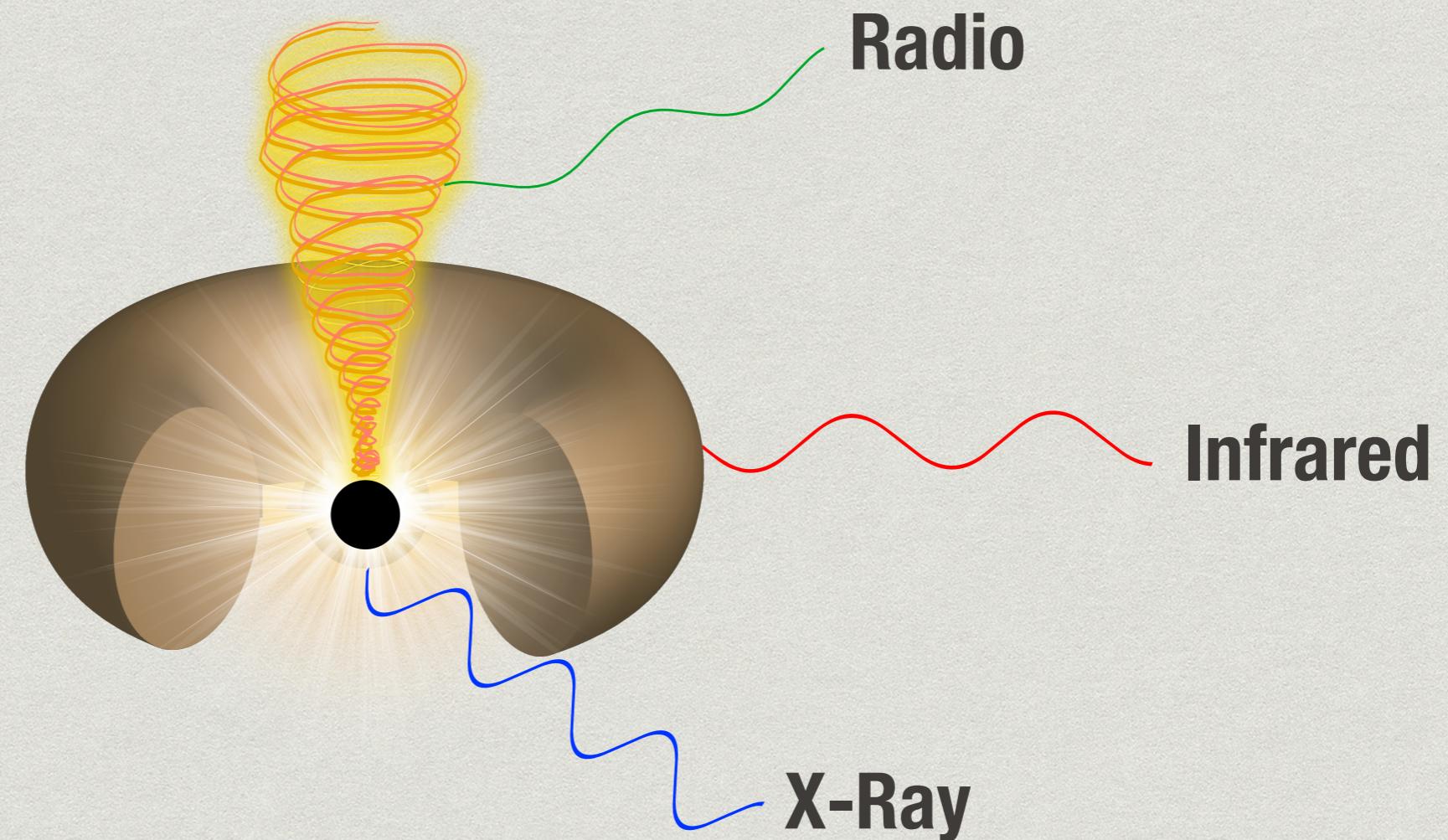


Chandra

# GOAL:

## 1) IDENTIFY A POPULATION OF GALAXIES HOSTING AN ACTIVE BLACK HOLE OVER A LARGE CHUNK OF COSMIC TIME

Requires a  
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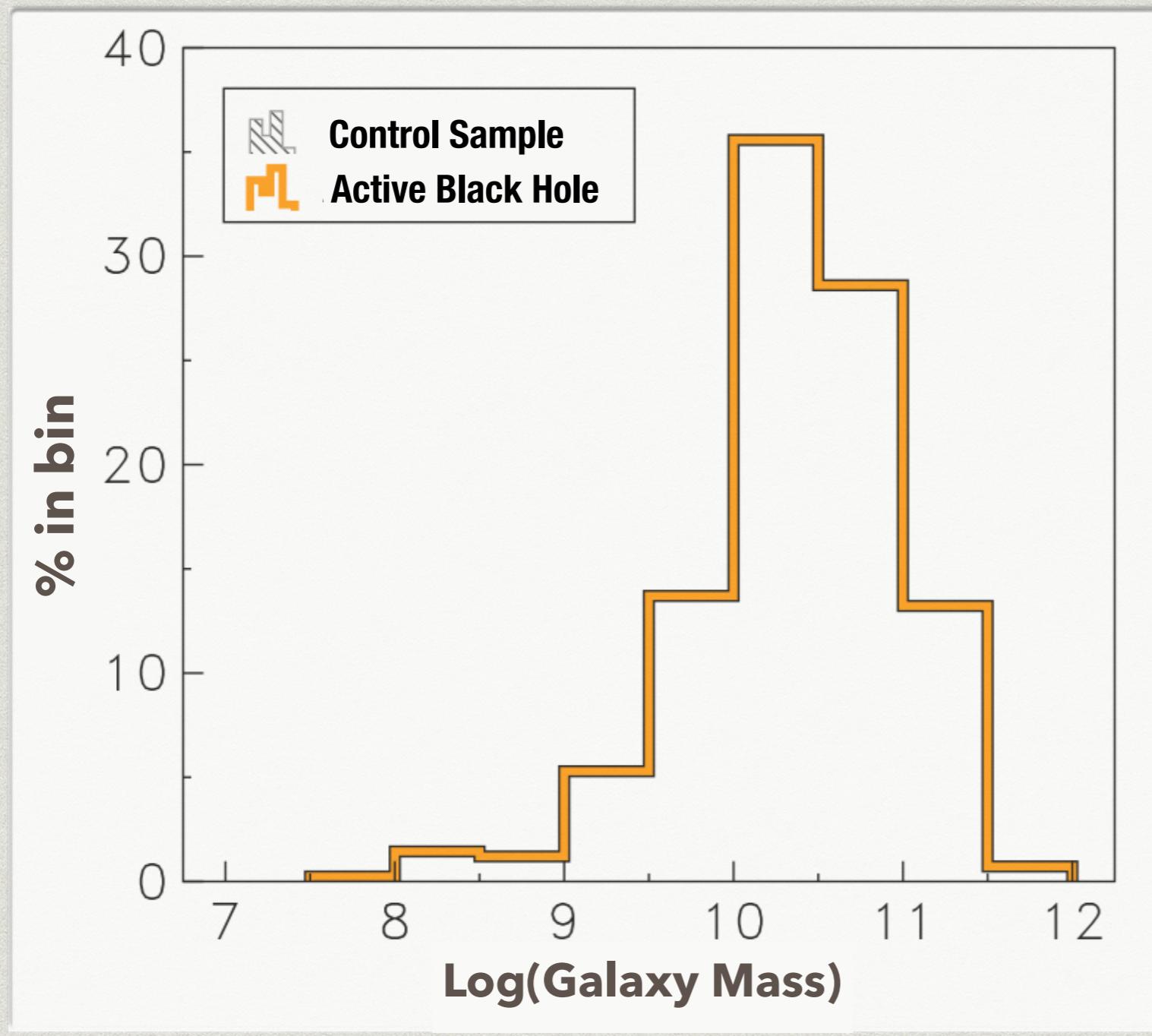


## Control Sample

Image Credit: NASA/JPL-Caltech

# GOAL:

## 2) COMPARE THE STAR-FORMATION ACTIVITY OF GALAXIES HOSTING AN ACTIVE BLACK HOLE TO THOSE WITHOUT

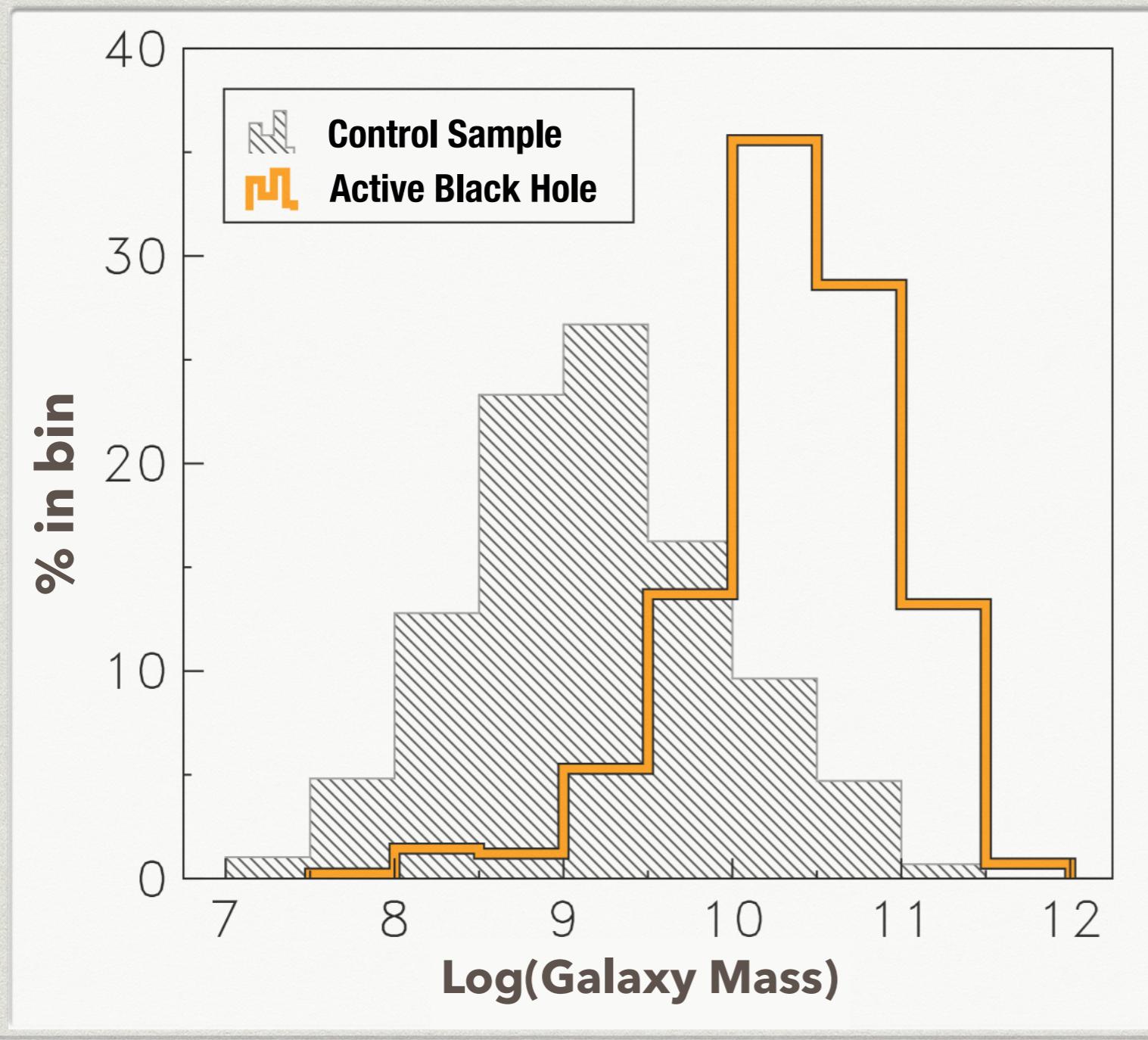


**Active supermassive black holes are preferentially hosted in galaxies with a high stellar mass (e.g., Aird+12)**

**A galaxy's stellar mass is tightly correlated with its star-formation rate (e.g., Noeske+07)**

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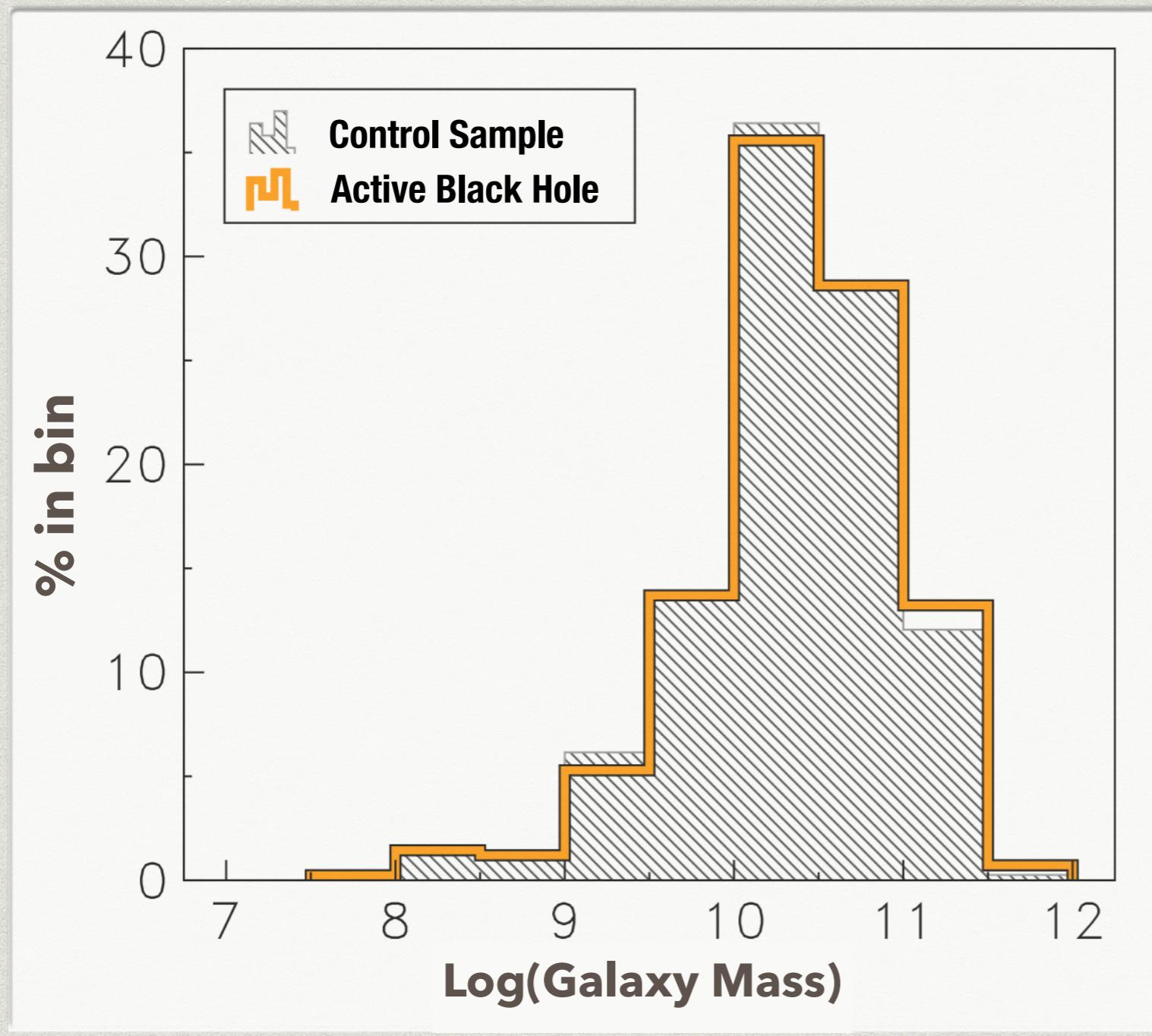


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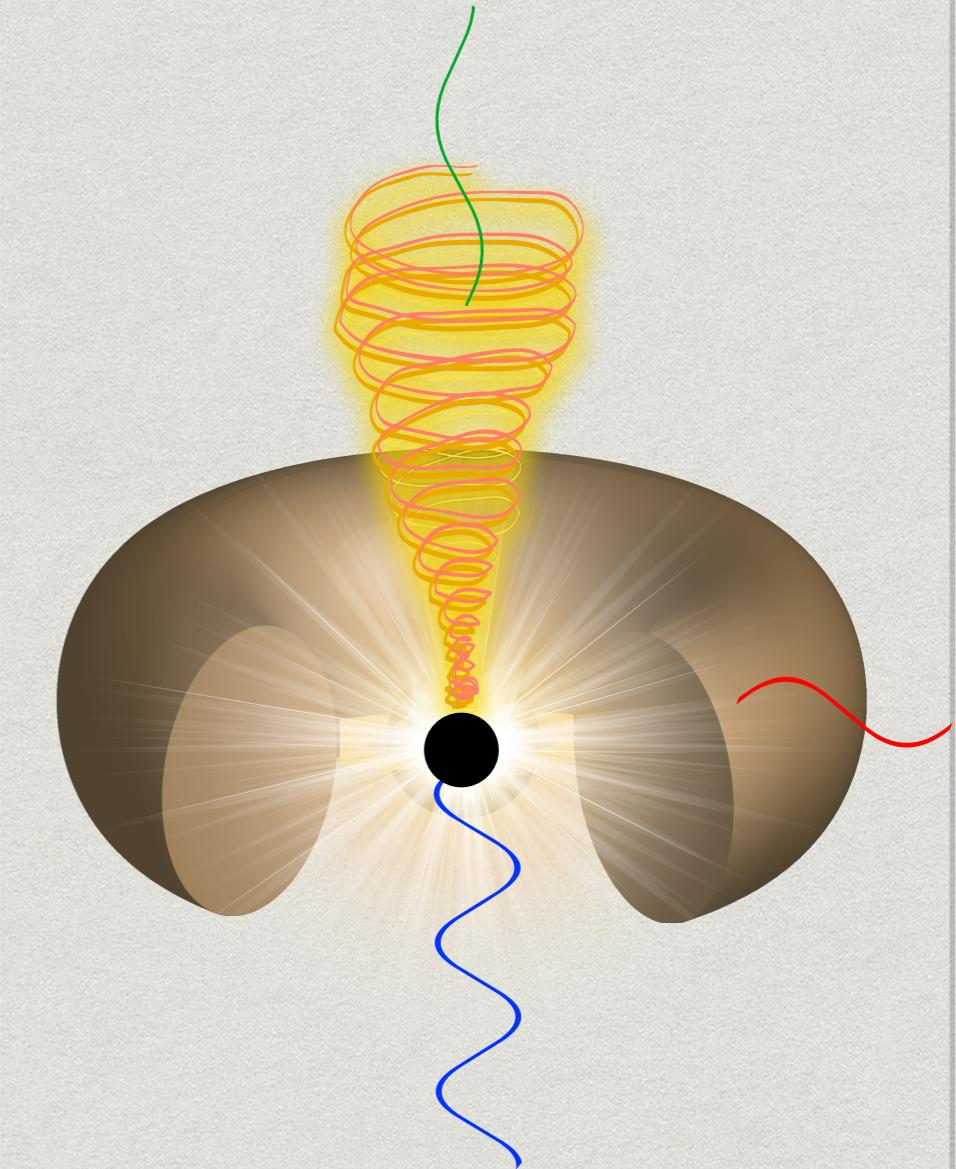
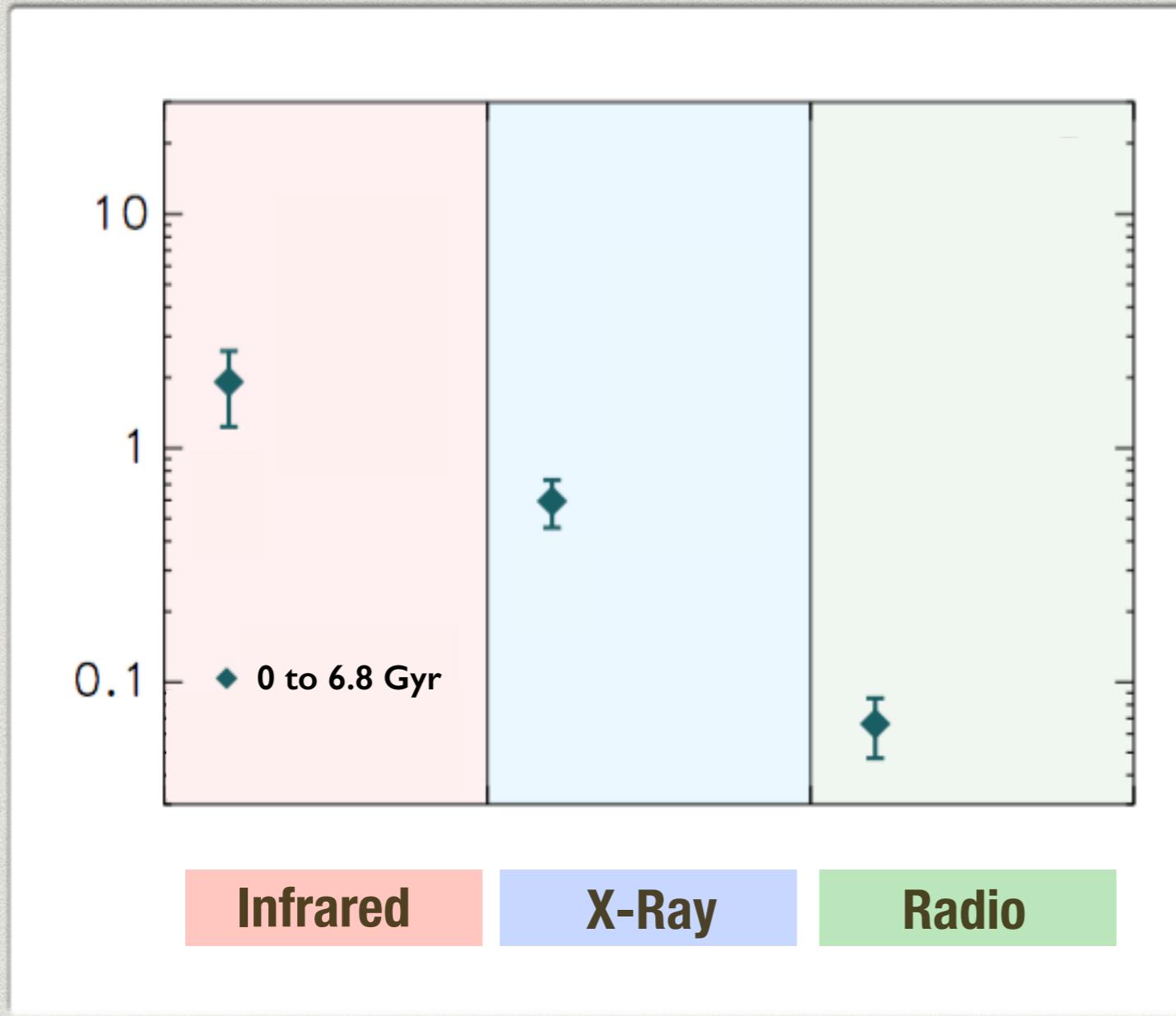


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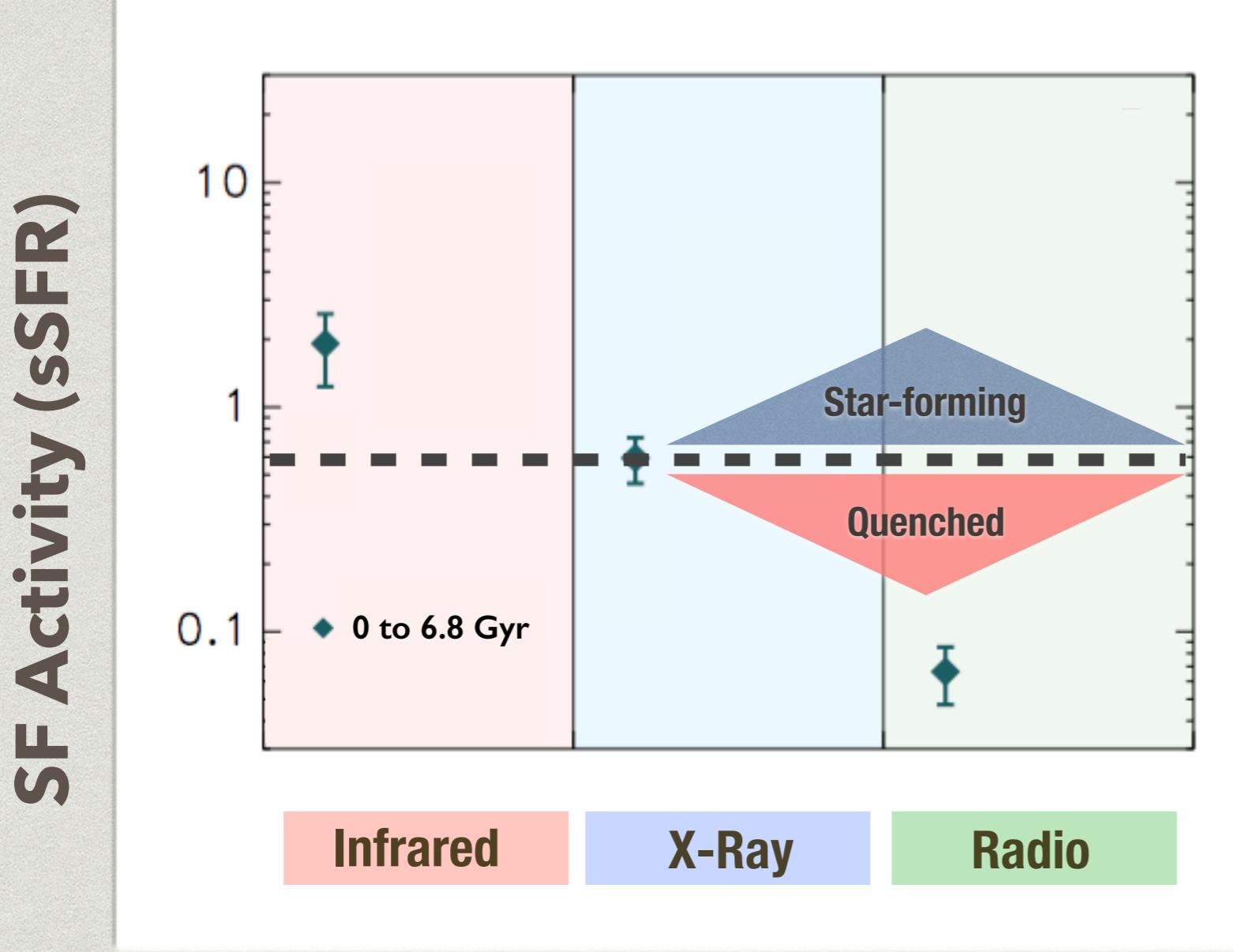
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# GOAL: COMPARE THE STAR-FORMATION ACTIVITY OF GALAXIES HOSTING AN ACTIVE BLACK HOLE TO THOSE WITHOUT

SF Activity (sSFR)



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**Infrared w/ Active Black Hole**

**Star forming hosts**

**X-ray w/ Active Black Hole**

**Straddles between star-forming and quiescent**

**Radio w/ Active Black Hole**

**Quenched hosts**

# GOAL: COMPARE THE STAR-FORMATION ACTIVITY OF GALAXIES HOSTING AN ACTIVE BLACK HOLE TO THOSE WITHOUT

Infrared w/ Active Black Hole



Younger Galaxies

X-Ray w/ Active Black Hole



Transitional

Radio w/ Active Black Hole

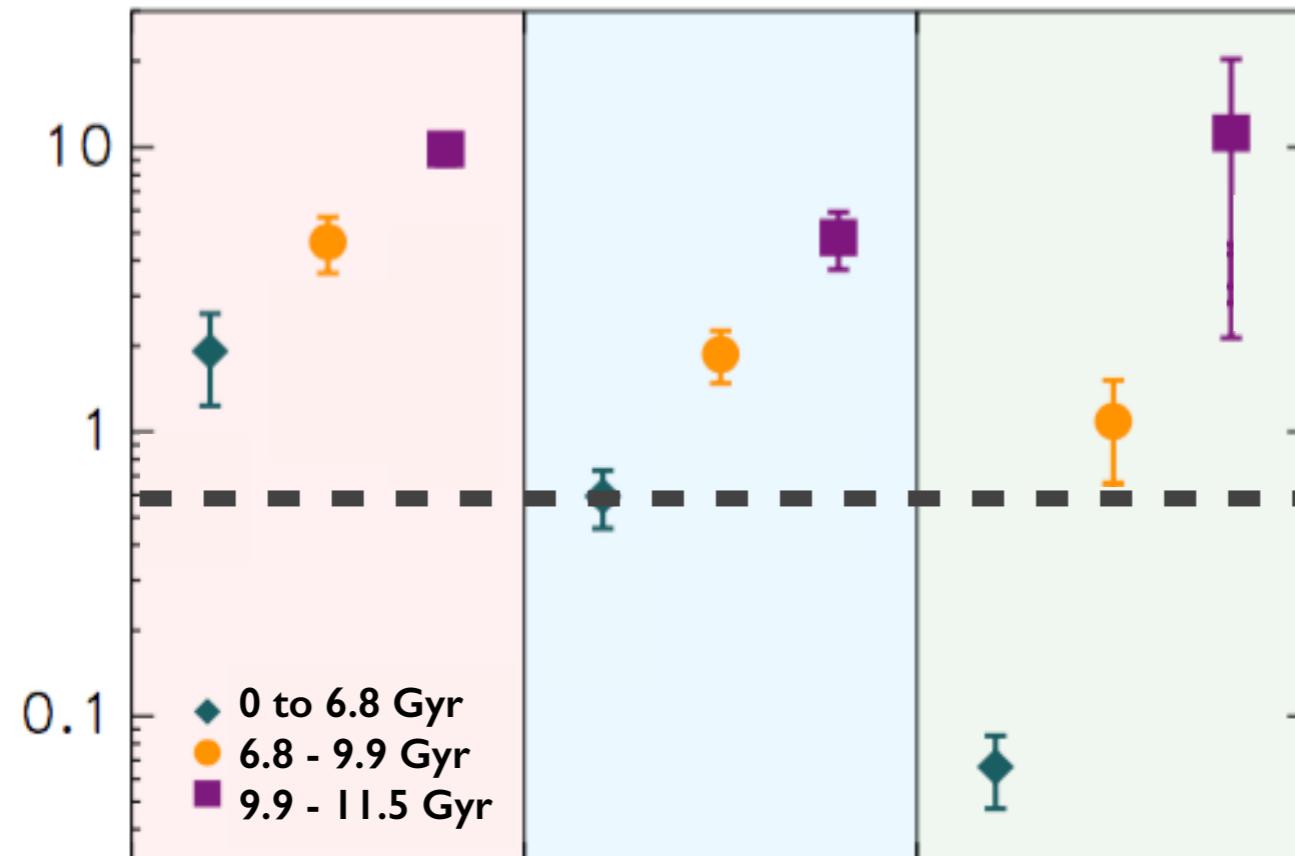


Older Galaxies

Image Credit: NASA/JPL-Caltech

# GOAL: COMPARE THE STAR-FORMATION ACTIVITY OF GALAXIES HOSTING AN ACTIVE BLACK HOLE TO THOSE WITHOUT

SF Activity (sSFR)



Infrared

X-Ray

Radio

Infrared w/ Active Black Hole

Star forming hosts

X-ray w/ Active Black Hole

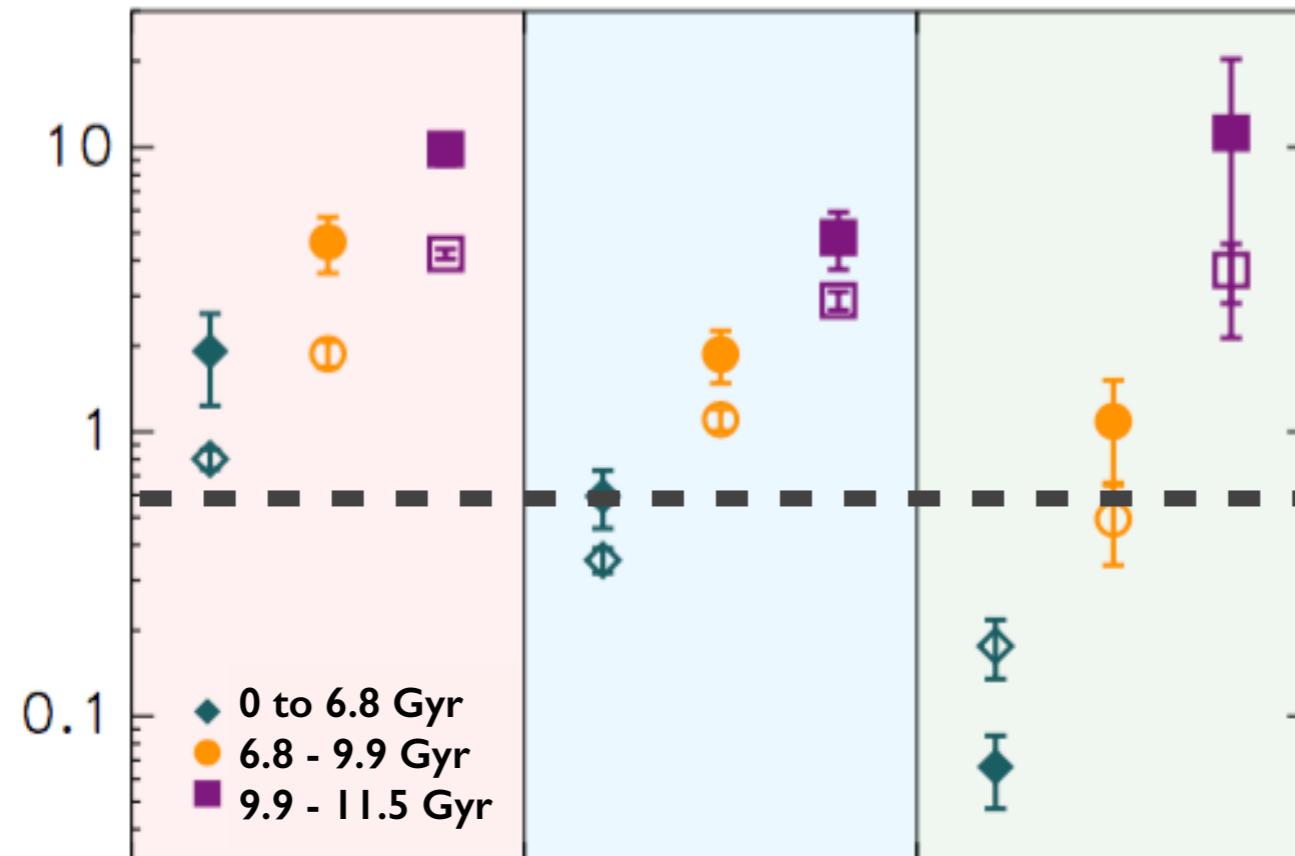
Star forming hosts

Radio w/ Active Black Hole

Star forming hosts

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X-Ray

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Infrared w/ Active Black Hole

Star forming hosts

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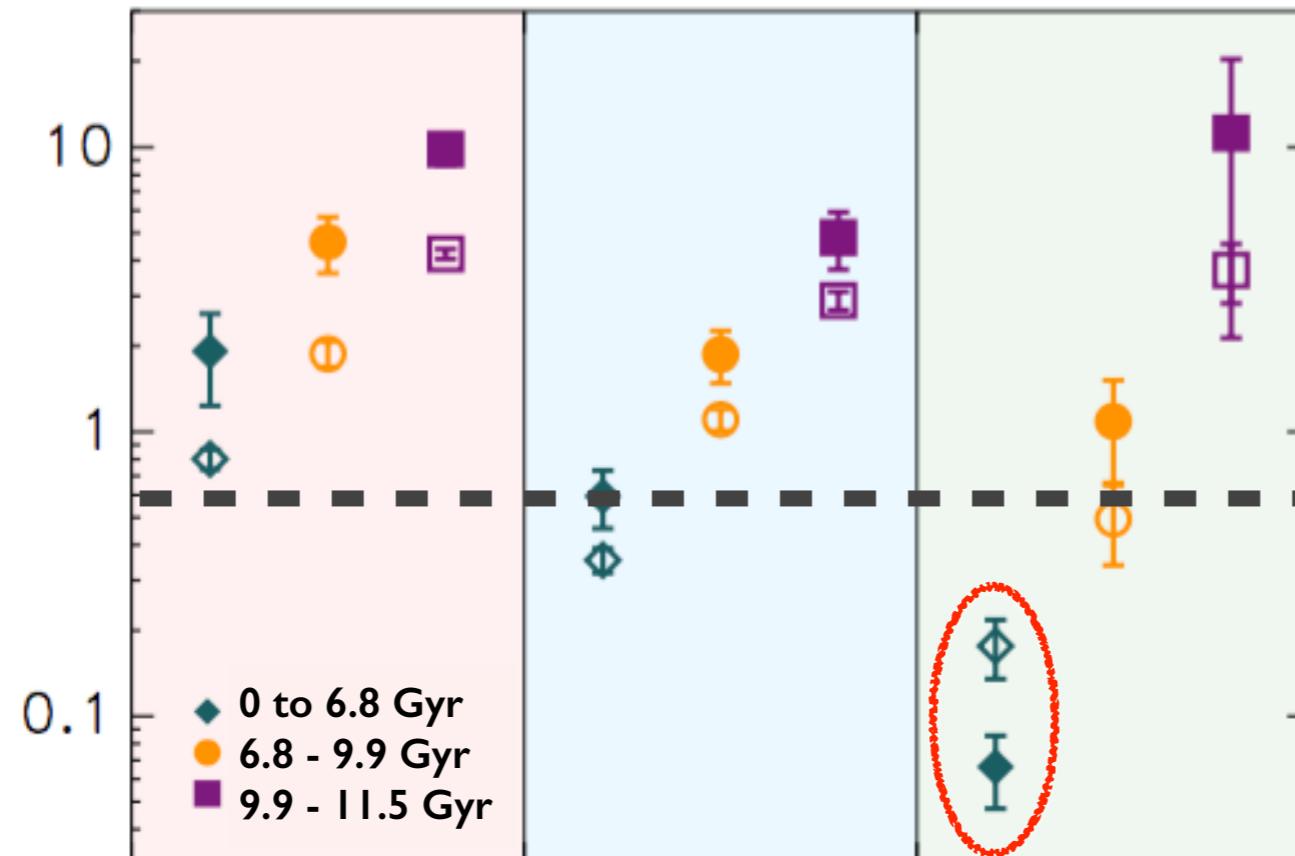
Star forming hosts

“Normal” Galaxies

Lower SF activity!

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SF Activity (sSFR)



Infrared w/ Active Black Hole

Star forming hosts

X-ray w/ Active Black Hole

Star forming hosts

Radio w/ Active Black Hole

Star forming hosts

“Normal” Galaxies

Lower SF activity!

# SUMMARY

- The mean star-formation activity of the control sample or “normal” galaxies tends to be lower than galaxies hosting an active supermassive black hole
- This calls into question the significance of black hole feedback being a primary mechanism for the suppression of star-formation in galaxies
- The only instance the above is not true is the radio-selected galaxies in the “local” universe

... what about the rest of cosmic time?